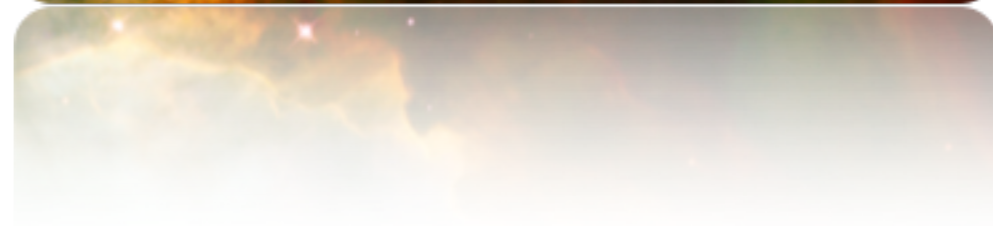


# The Meudon PDR code

Astrosim School

July 2017



## Fortran code

- developed over 30 years
- several versions
- public access  
[www.ism.obspm.fr](http://www.ism.obspm.fr)
- services in progress

## Programmers

Jacques Le Bourlot  
Evelyne Roueff  
Franck Le Petit  
Emeric Bron  
Benjamin Godard

## Goals

- chemical structure of interstellar clouds
- interactions with photons and cosmic rays
- treat detailed microphysical processes
- analyse their couplings
- interpret observations



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

1. Introduction on PDRs
2. Assumptions
3. User Guide - first steps
4. Examples



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1. Introduction on PDRs
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3. User Guide - first steps
4. Examples
5. Physics & algorithms
6. Applications
7. Recent updates
8. Conclusions



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## PDR

- ☐ Photon Dominated Regions  
or
- ☐ Photodissociation Regions

⇒ Interstellar environment where the FUV radiation impacts the chemical / thermal state and evolution of the gas and dust

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## FUV photons

- 6 eV — 13.6 eV

## impacts

- heating source
  - dust - IR emission
  - gas - photoelectric effect
- photodissociation / ionization
- excitation / pumping

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## FUV photons

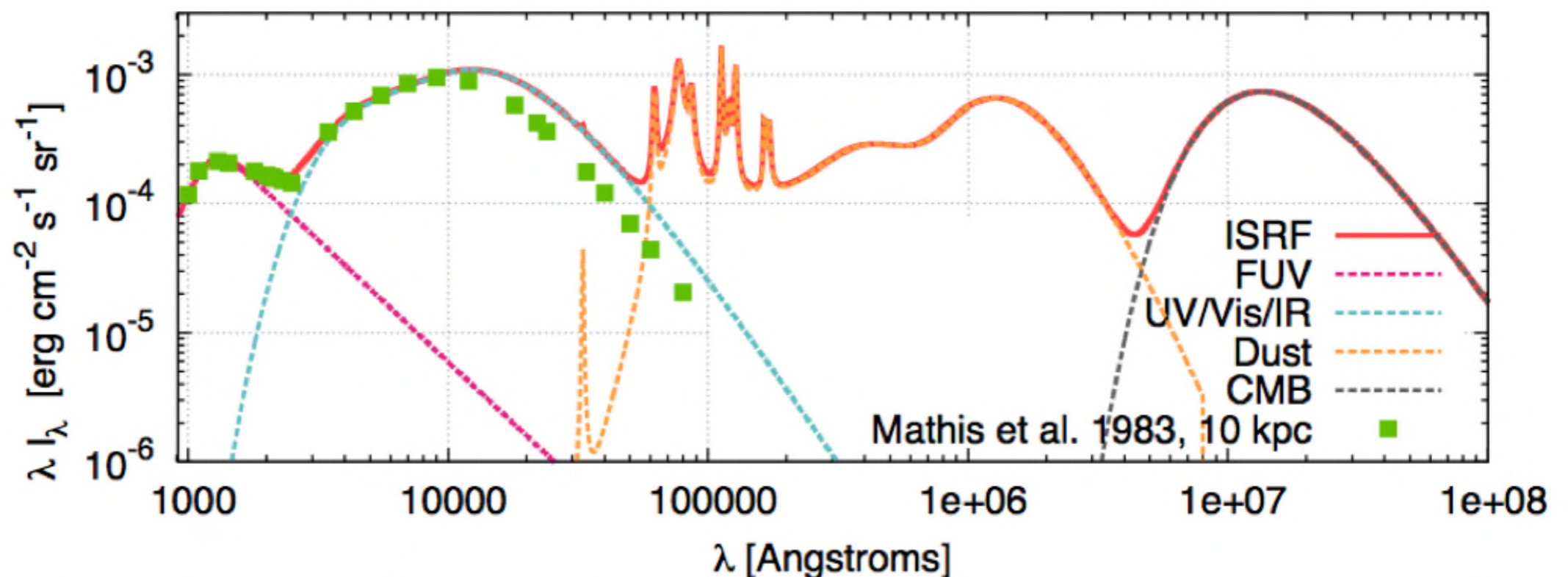
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## FUV photons

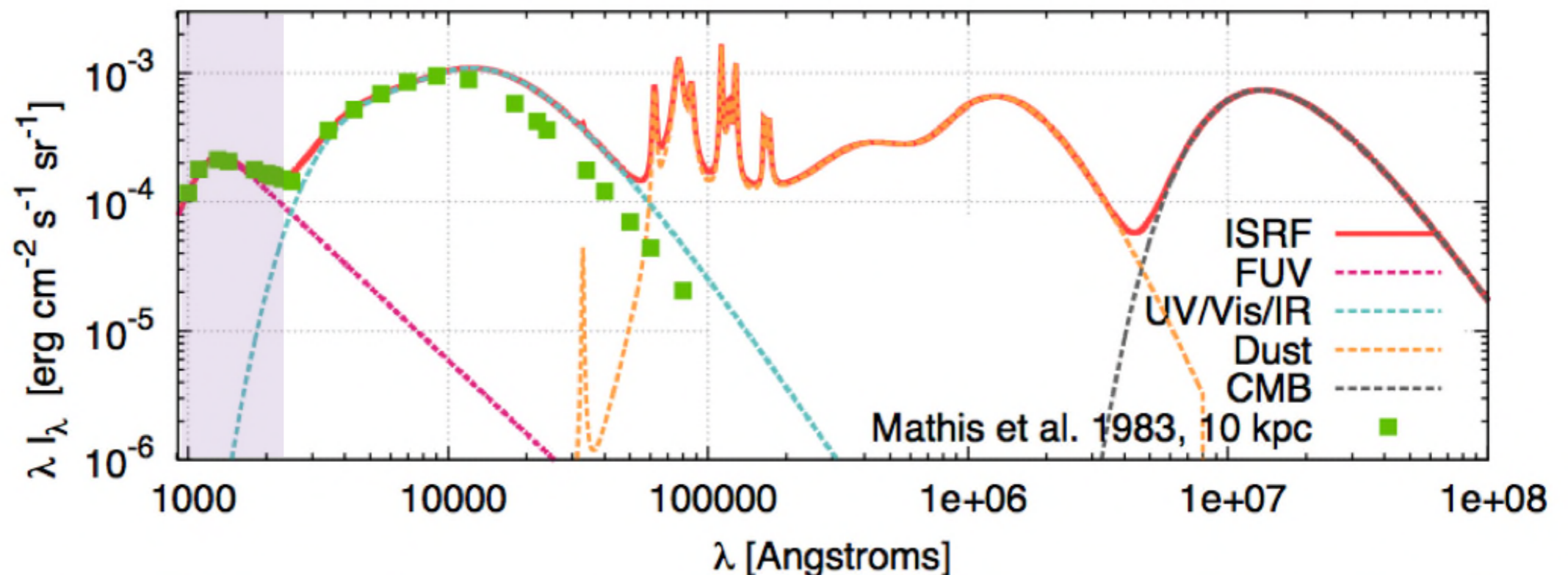
- 6 eV — 13.6 eV
- ISRF - Mathis et al. (1983)  
flux:  $1.6 \times 10^{-3} \text{ erg cm}^{-2} \text{ s}^{-1}$
- often parametrized with  $G_0$

## impacts

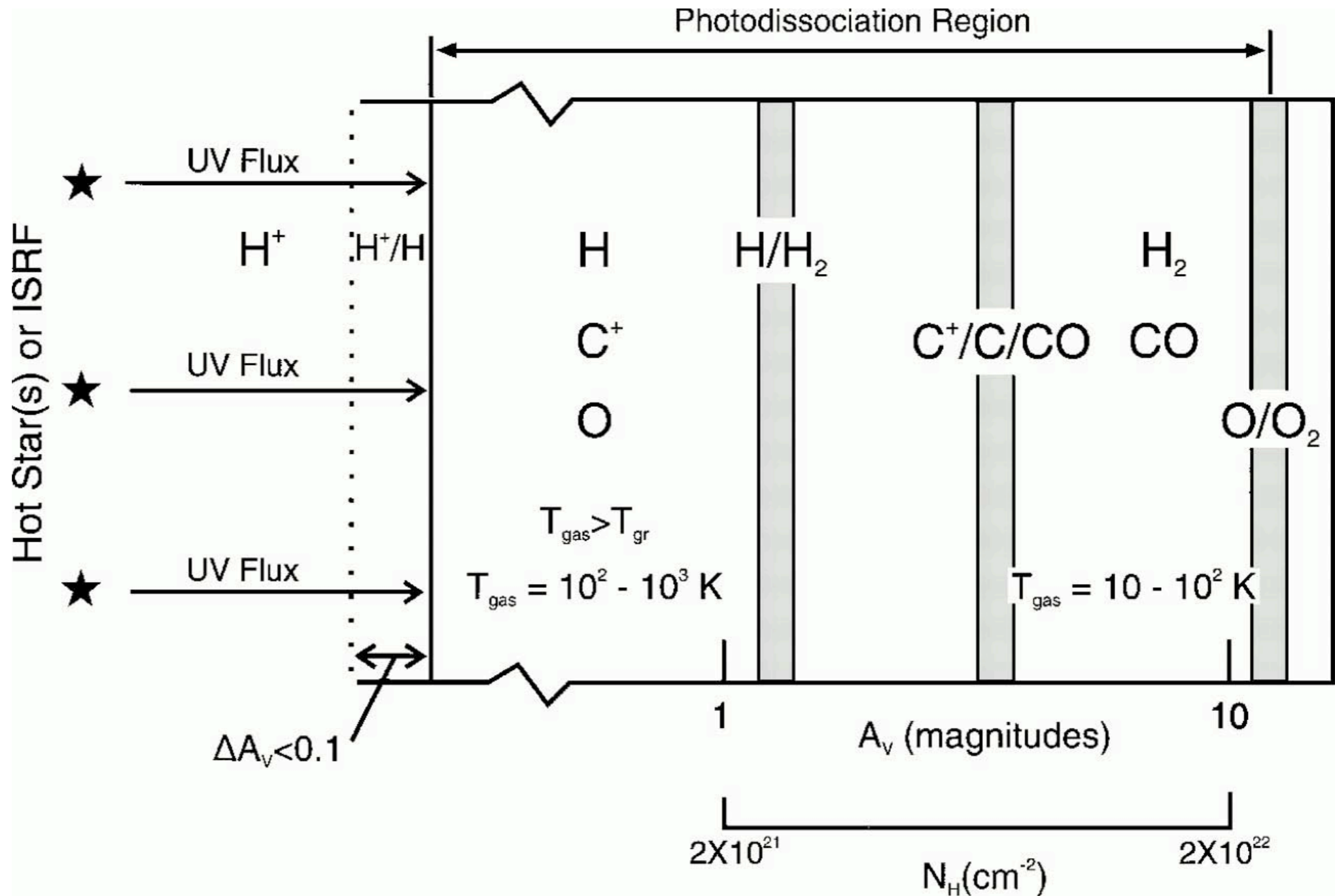
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## Structure of a PDR



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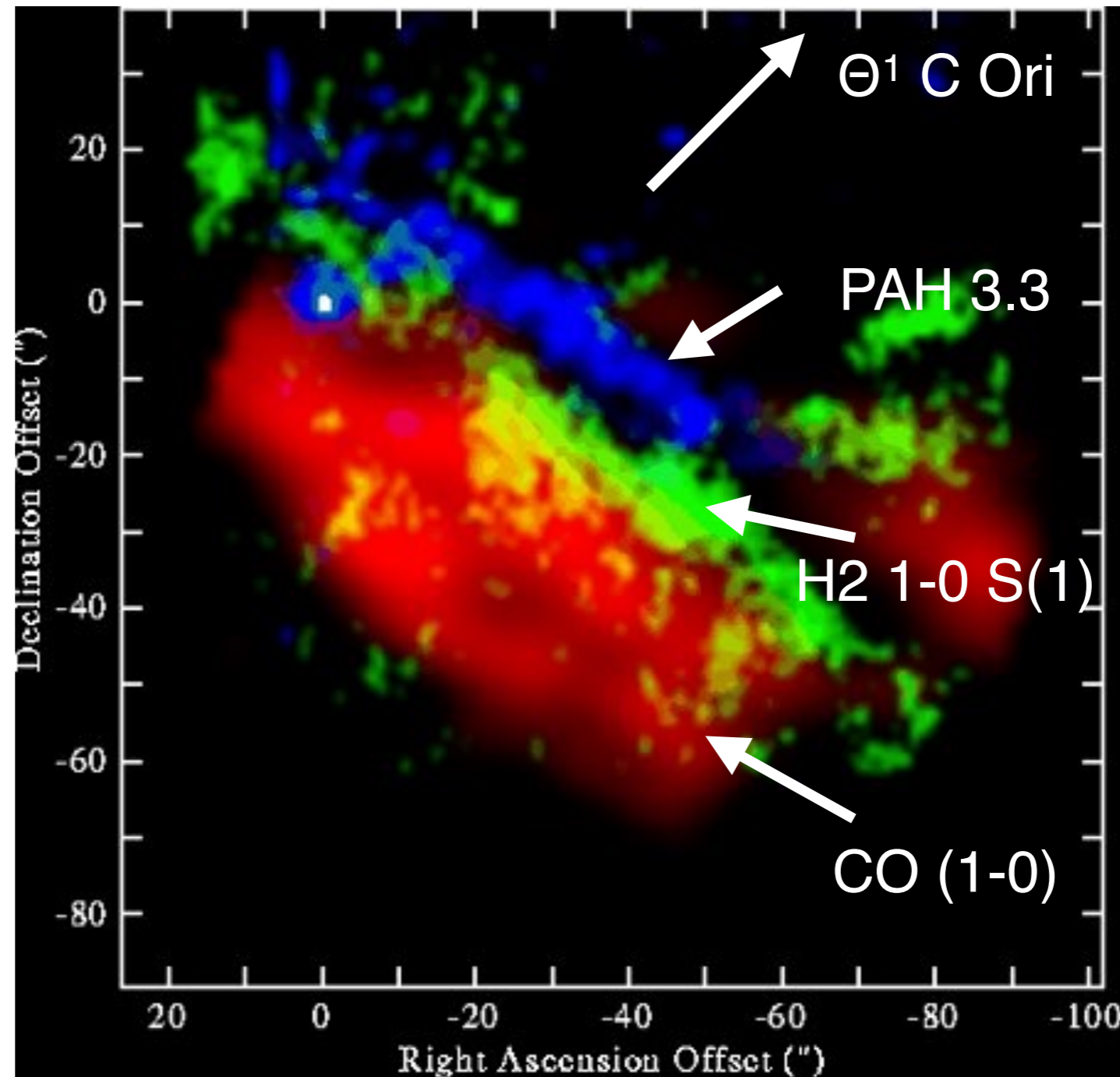
## Example



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## Example



Tielens, Meixner, et al. (1993)

## Outline

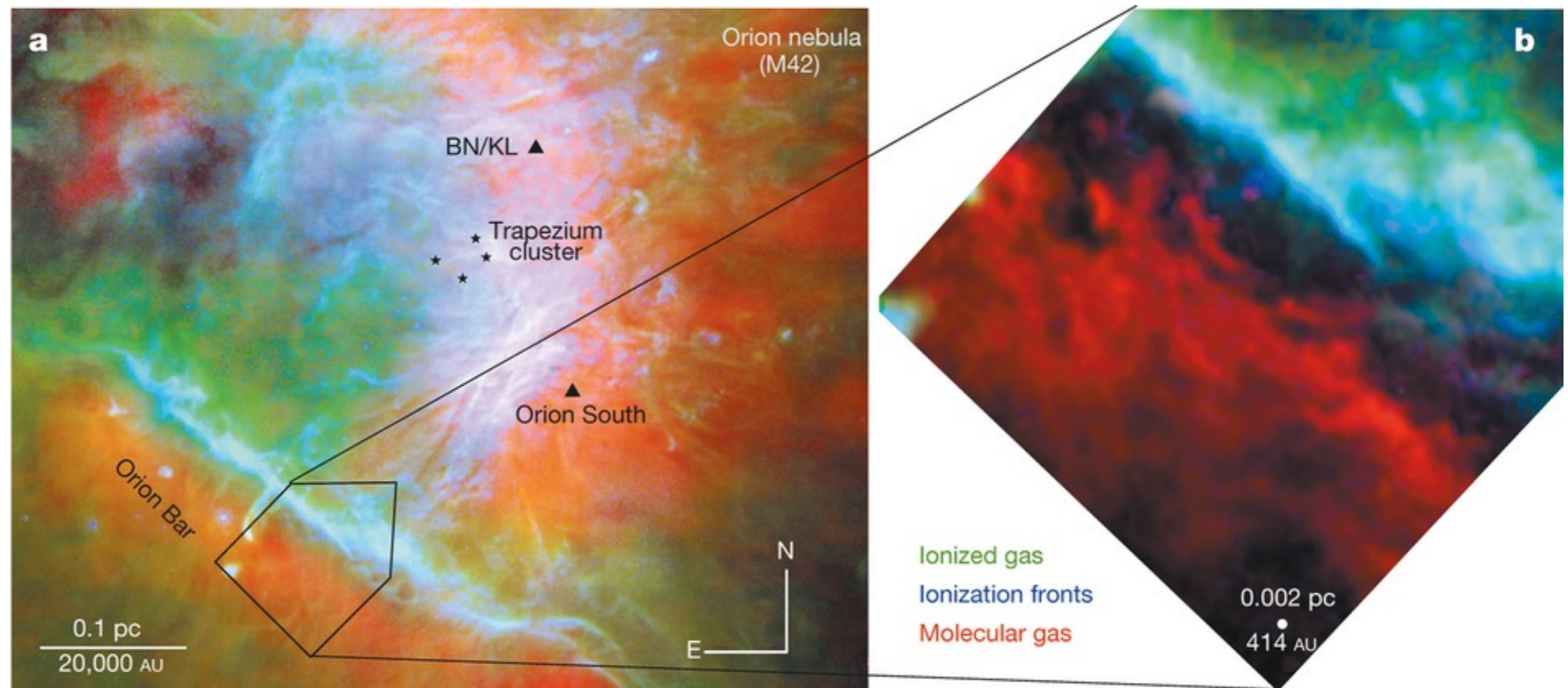
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Goicoechea et al. (2016)

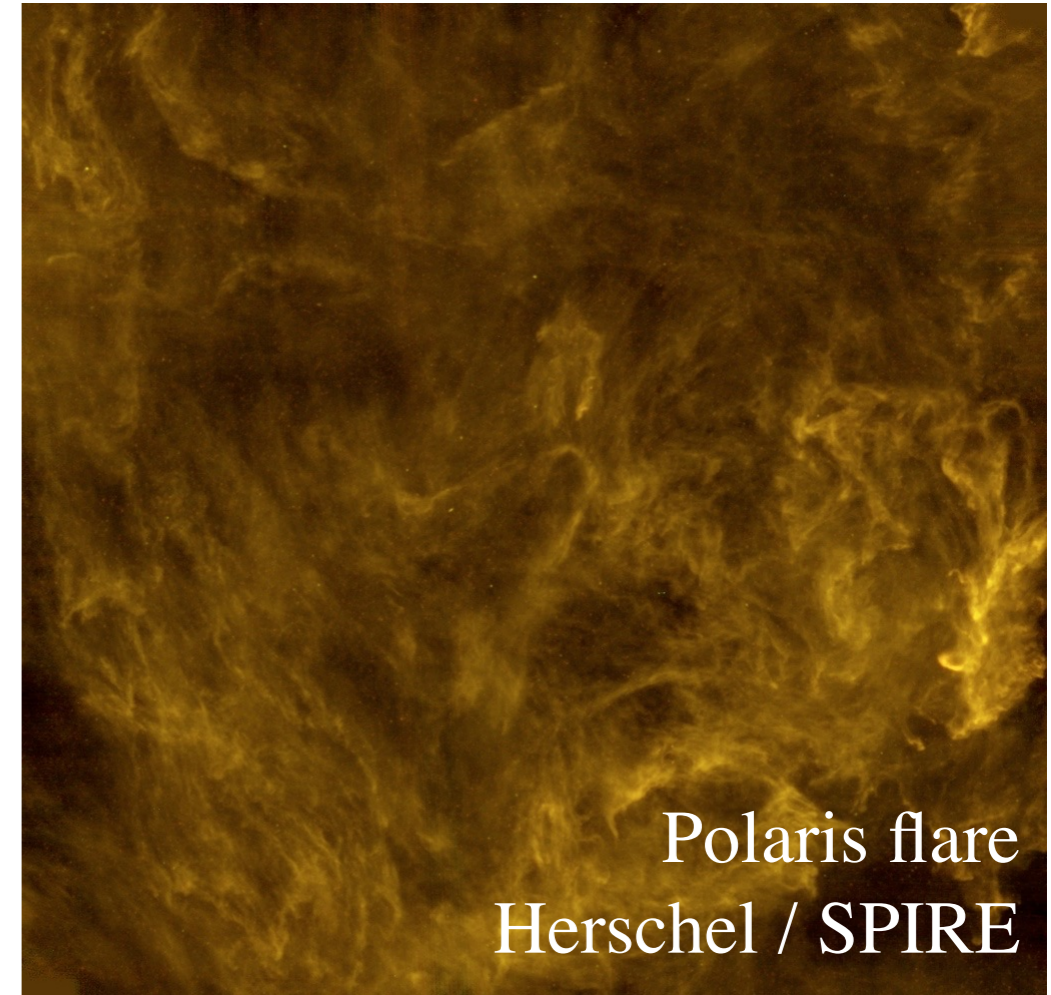


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- Warm Neutral Medium
- Diffuse clouds
- star forming regions
- Reflection nebulae
- Neutral gas around PN
- Photodiss winds from red giants and AGB
- Protoplanetary disks

## Examples

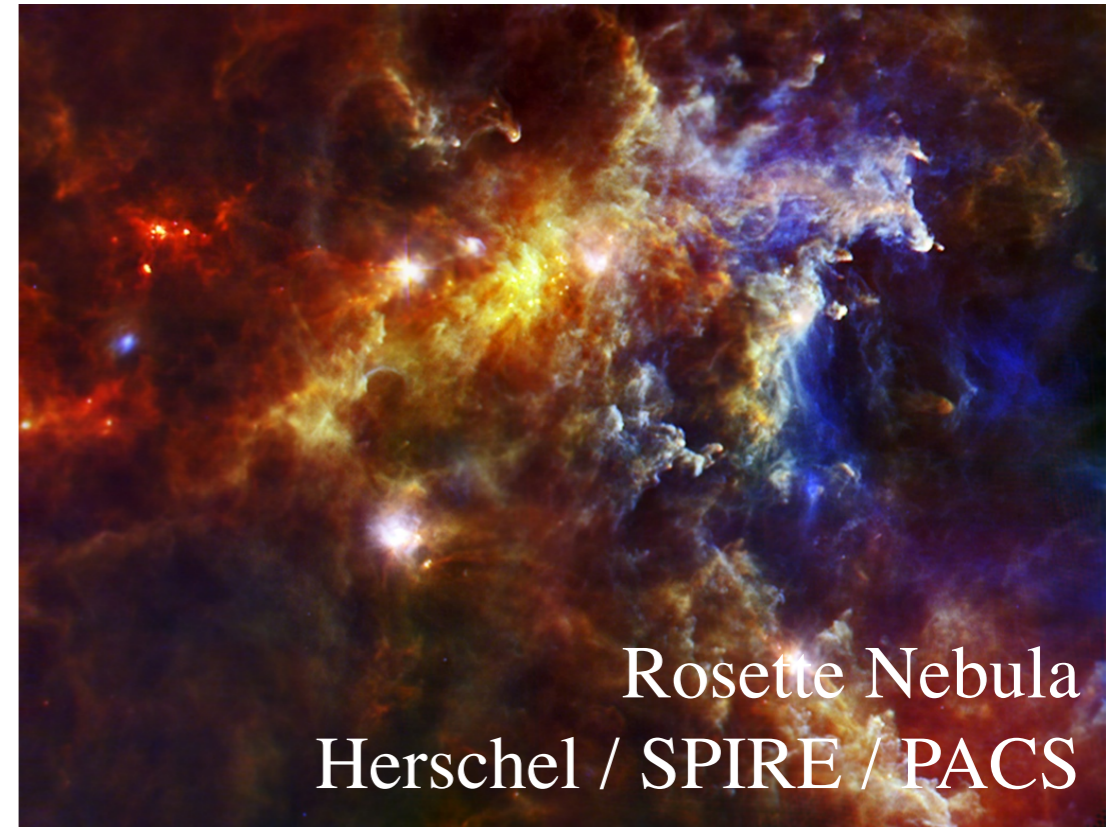


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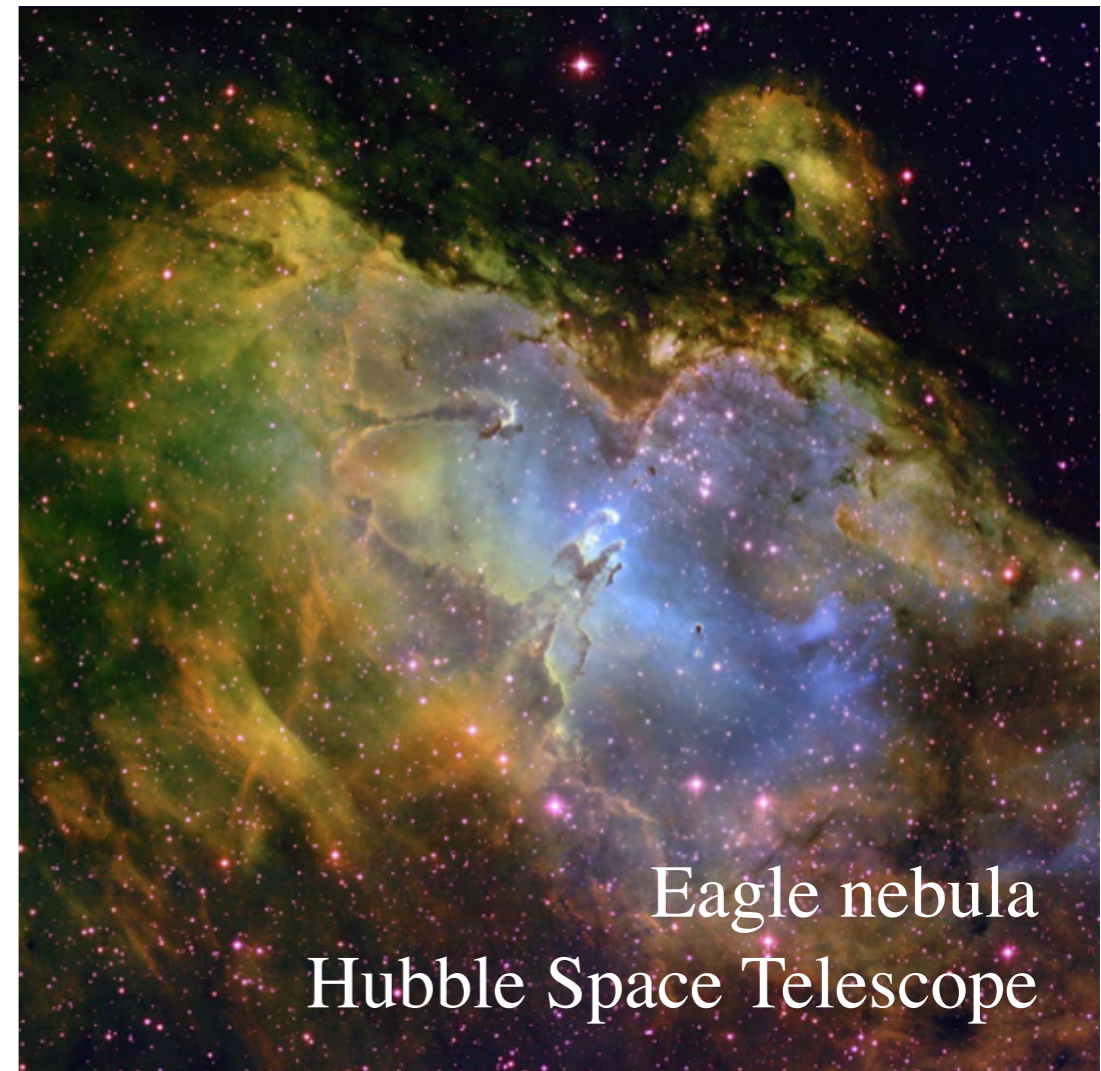


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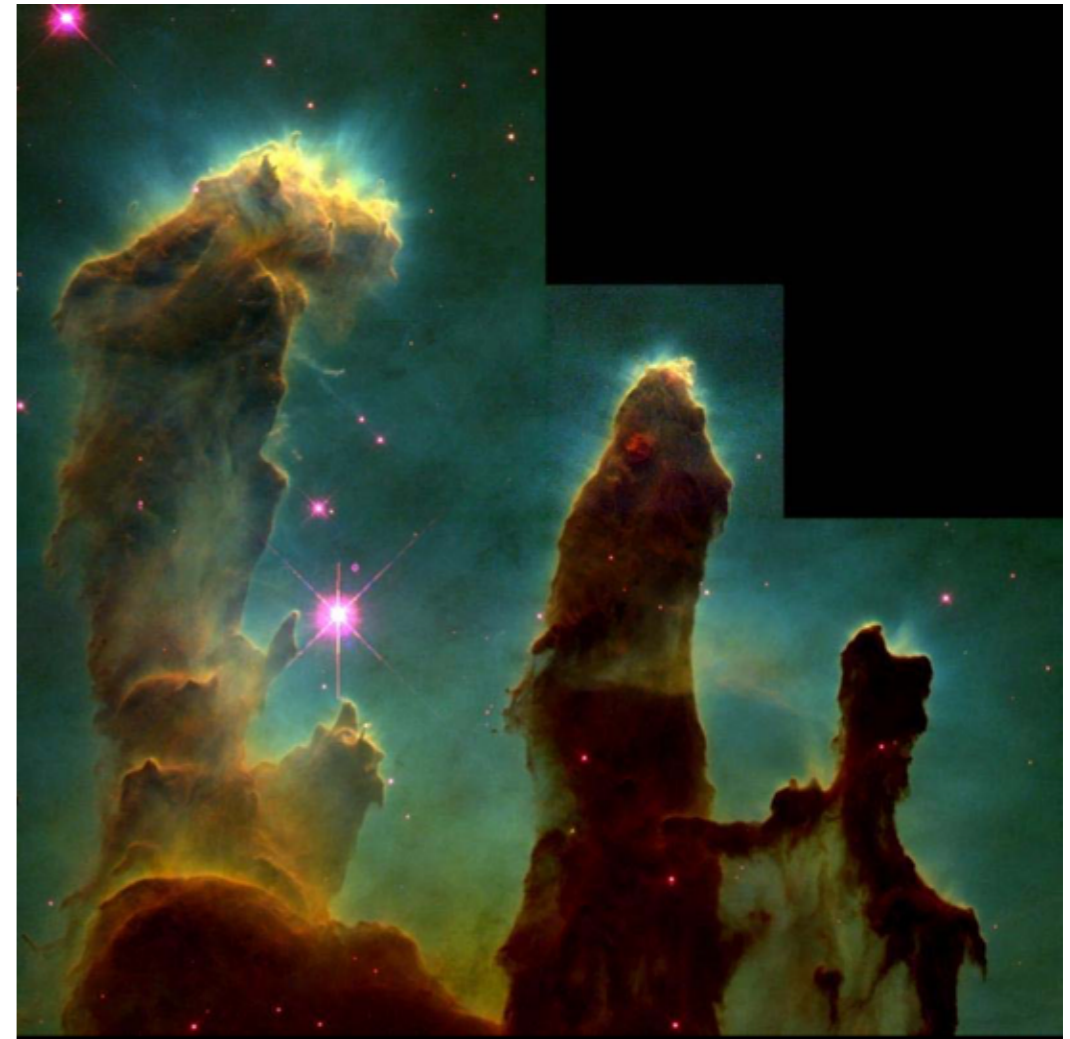


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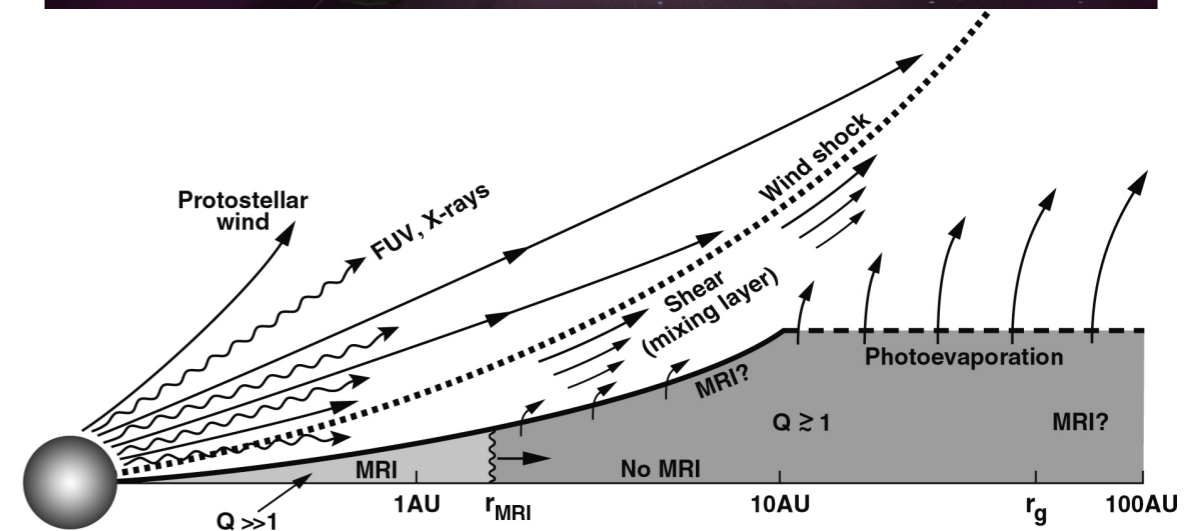
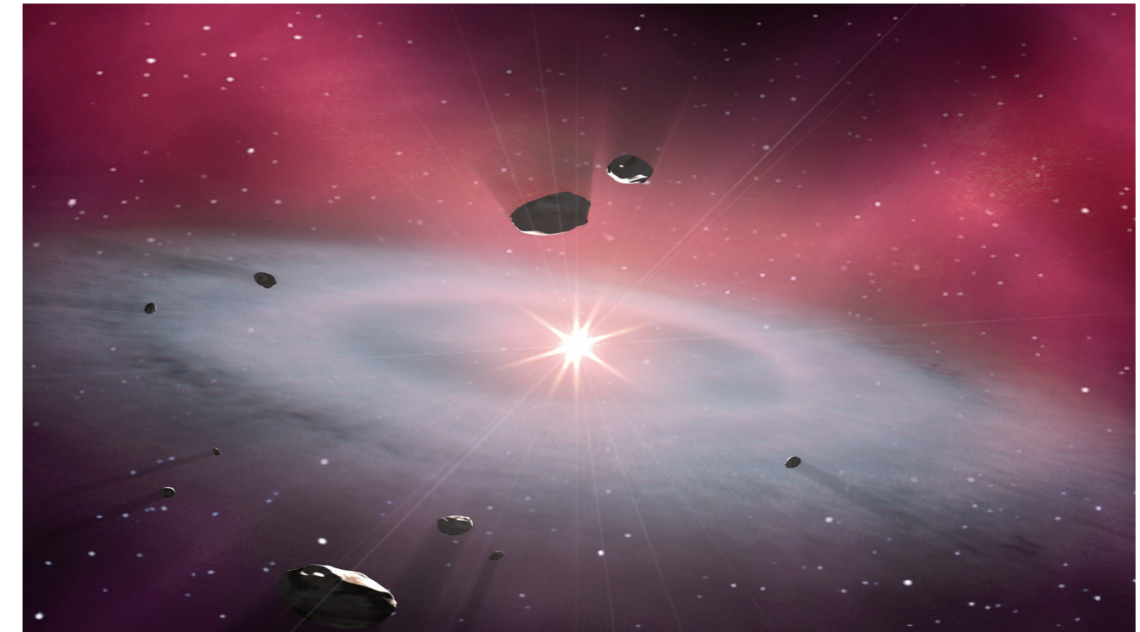
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## Why studying PDRs ?

- most of interstellar matter in PDRs
- physics of PDRs
  - ✓ thermal instability → CNM / WNM
  - ✓ ionization → coupling with B
  - ✓ dust charge → dynamics
- emission of PDR
  - ✓ PAH & grain → large fraction of IR photons
  - ✓ C<sup>+</sup>, O, H<sub>2</sub>, CO, ... → physical conditions
  - ✓ stratification → geometry of sources
- study of extragalactic environments

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chemical mixture  
mono fluid, no B

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$$\frac{Dn_s}{Dt} = \frac{\partial n_s}{\partial t} + (\mathbf{u} \cdot \nabla) n_s = -n_s \nabla \cdot \mathbf{u} + F_s - D_s - \nabla \cdot \frac{1}{m_s} \mathbf{j}_s$$

$$\frac{D\mathbf{u}}{Dt} = \frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\nabla V - \frac{1}{\rho} \nabla P + \frac{1}{\rho} \nabla \cdot \boldsymbol{\pi}$$

$$\frac{D\varepsilon}{Dt} = \frac{\partial \varepsilon}{\partial t} + (\mathbf{u} \cdot \nabla) \varepsilon = -\frac{P}{\rho} \nabla \cdot \mathbf{u} + \frac{1}{\rho} (\Gamma_{\text{vis}} + \Gamma - \Lambda) - \frac{1}{\rho} \nabla \cdot \mathbf{C}$$

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-- advection

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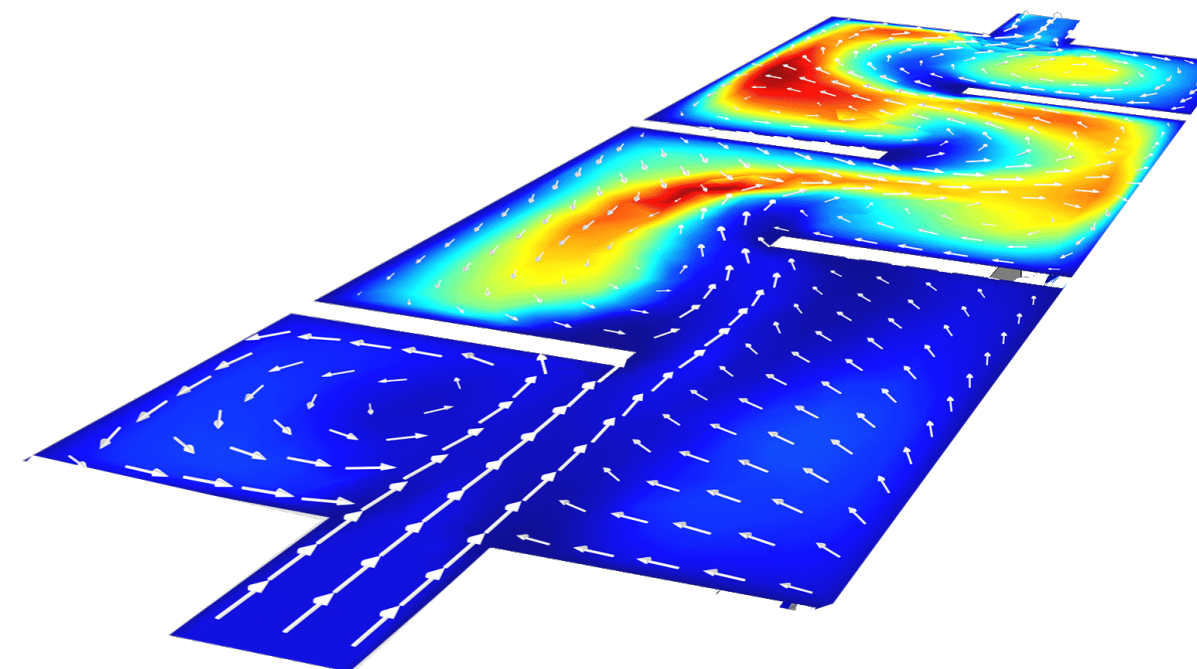
-- advection    -- compression    -- thermo-chemistry    -- diffusion

# Simplifications

In a perfect world  $\rightarrow$  3D, diffusive, with radiative transfer

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Main simplifications



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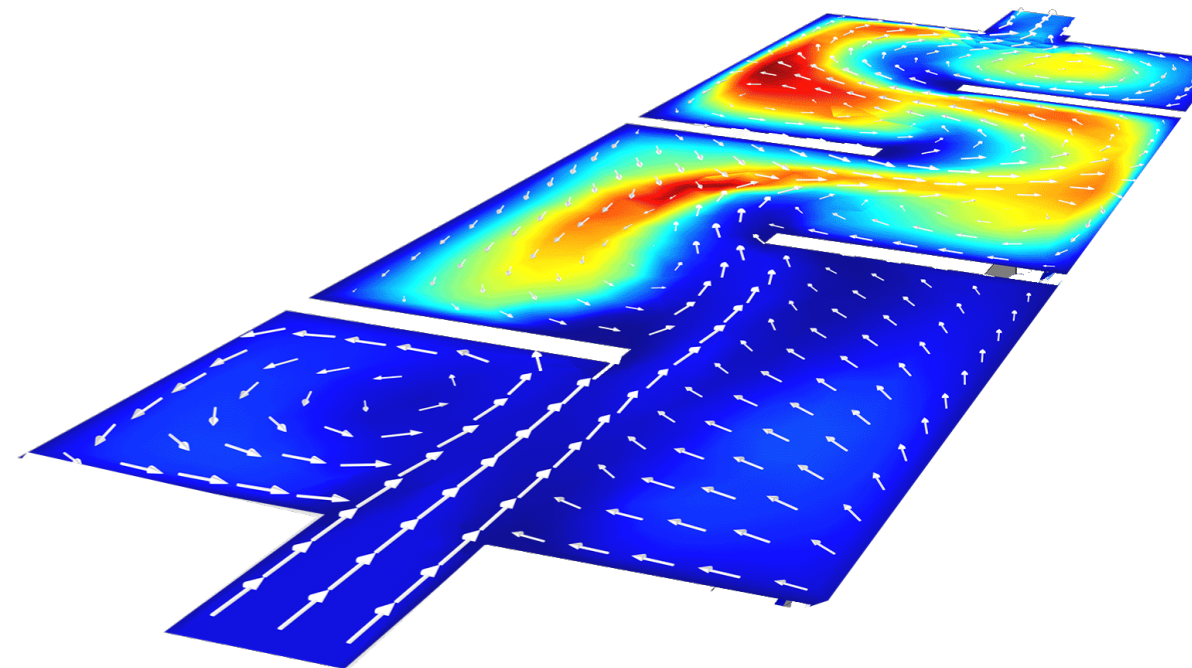
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Main simplifications

- number of spatial dimensions (2D, 1D, 0D)



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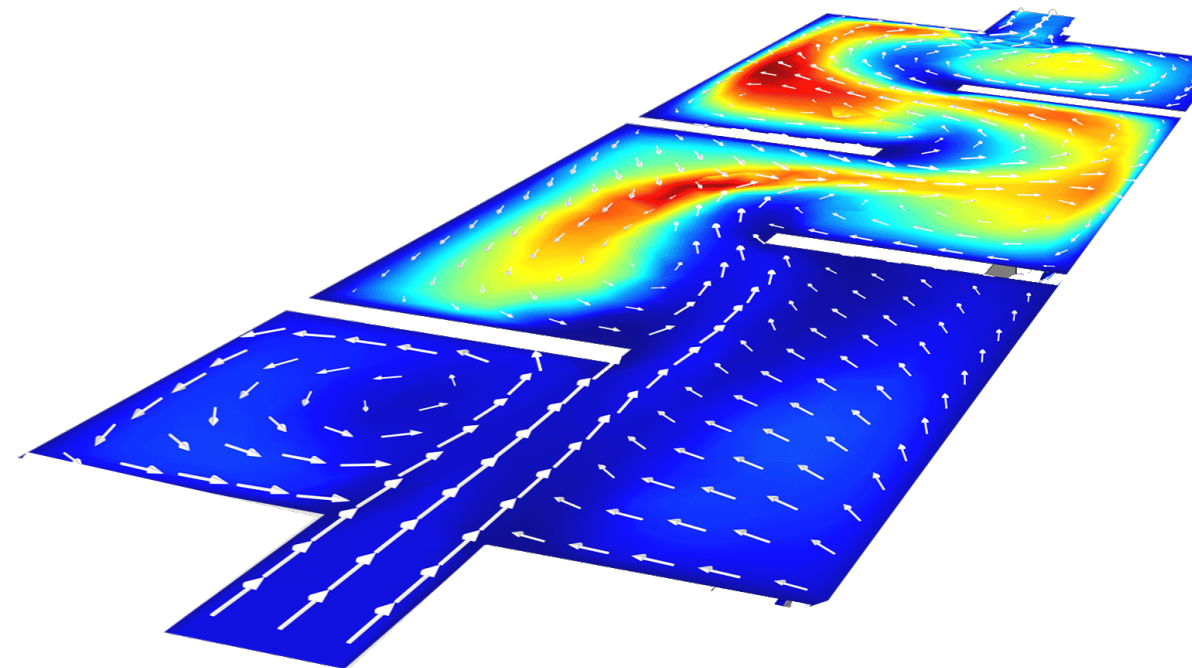
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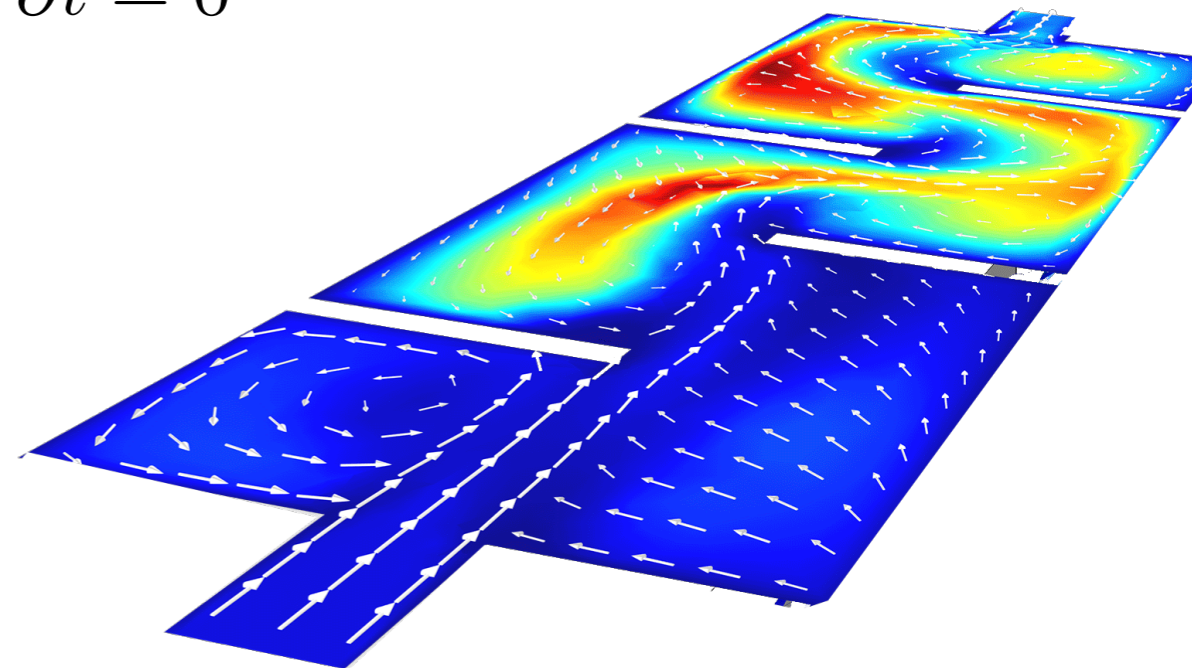
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- number of spatial dimensions (2D, 1D, 0D)
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- steady-state (stationary)  $\partial/\partial t = 0$



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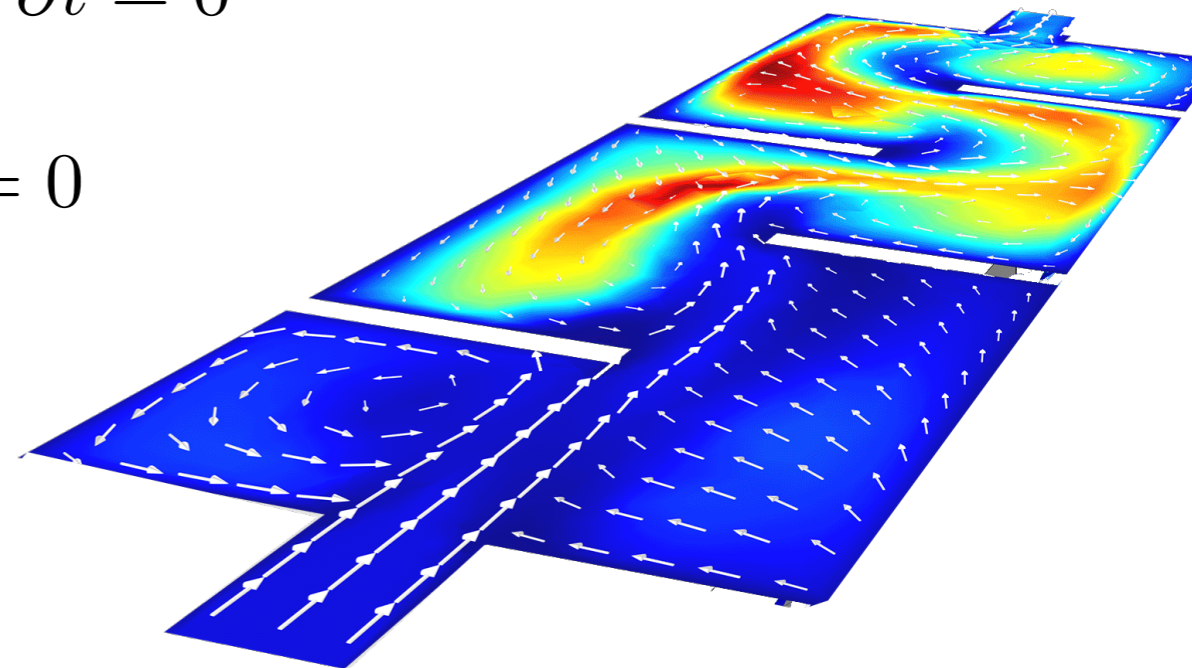
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Main simplifications

- number of spatial dimensions (2D, 1D, 0D)
- neglect diffusion
- steady-state (stationary)  $\partial/\partial t = 0$
- static (equilibrium)  $D/Dt = 0$

$$\triangle \frac{\partial}{\partial t} = 0 \neq \frac{D}{Dt} = 0$$



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# Numerical codes

## Outline

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	Nautilus	Meudon PXDR	CLOUDY	Paris-Durham shock	CHEMSES	RAMSES & KROME	
dimension	0	1	1	1	3	3	
dynamical		static	steady	✓	✓		
thermal		static	steady		✓		
chemical		static	steady		✓		
							free
	NAHOON	TDR	MAPPING	ENZO	ZEUS	NIRVANA	
dimension	0	1	1	3	3	3	
dynamical		✓	steady				✓
thermal		steady	steady	✓			✓
chemical		steady	steady				
							✓ diffusion

## The Meudon PDR code

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## The Meudon PDR code

- radiative transfer (UV to radio)  
gas and dust processes
- dust treatment  
charge, emissivities



### Outline

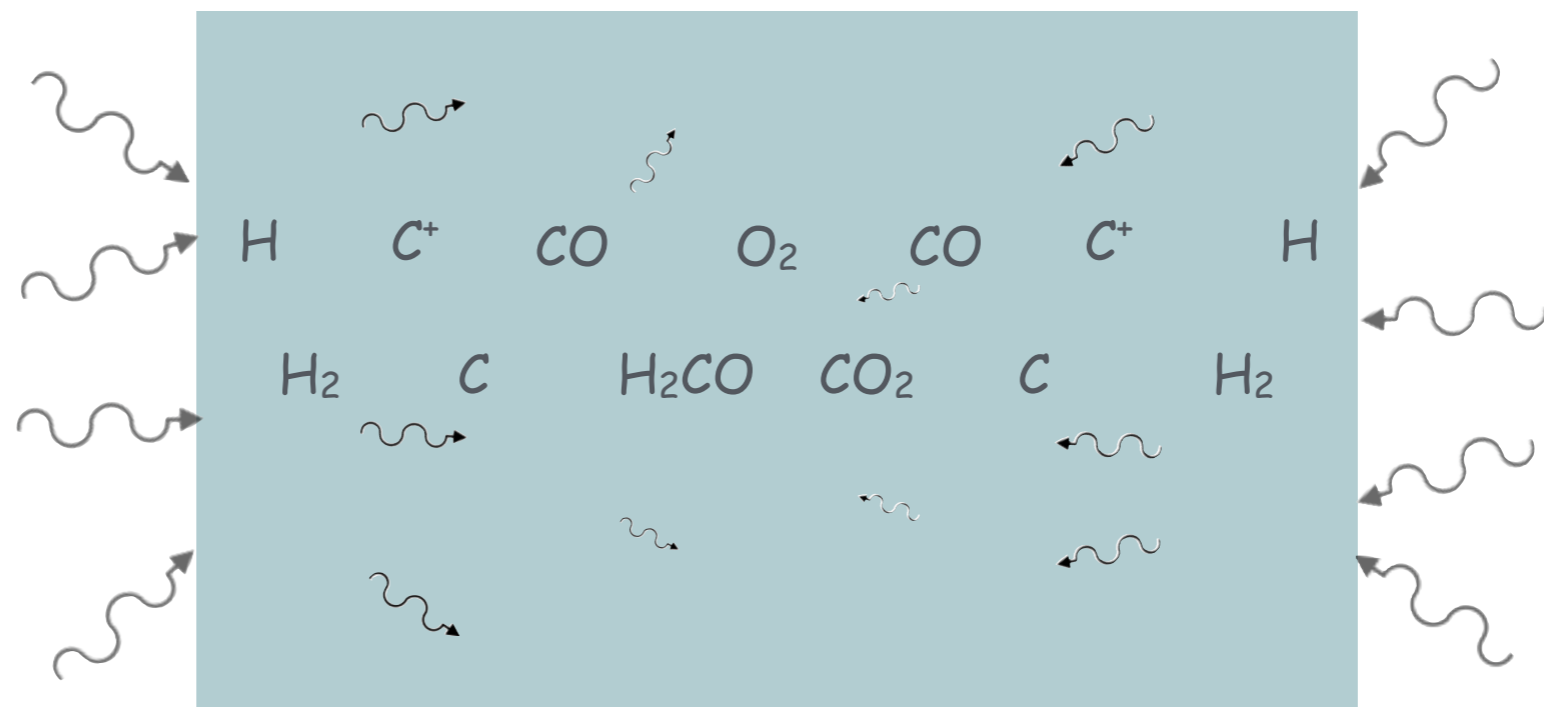
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gas and dust processes
- chemistry (hundreds species)
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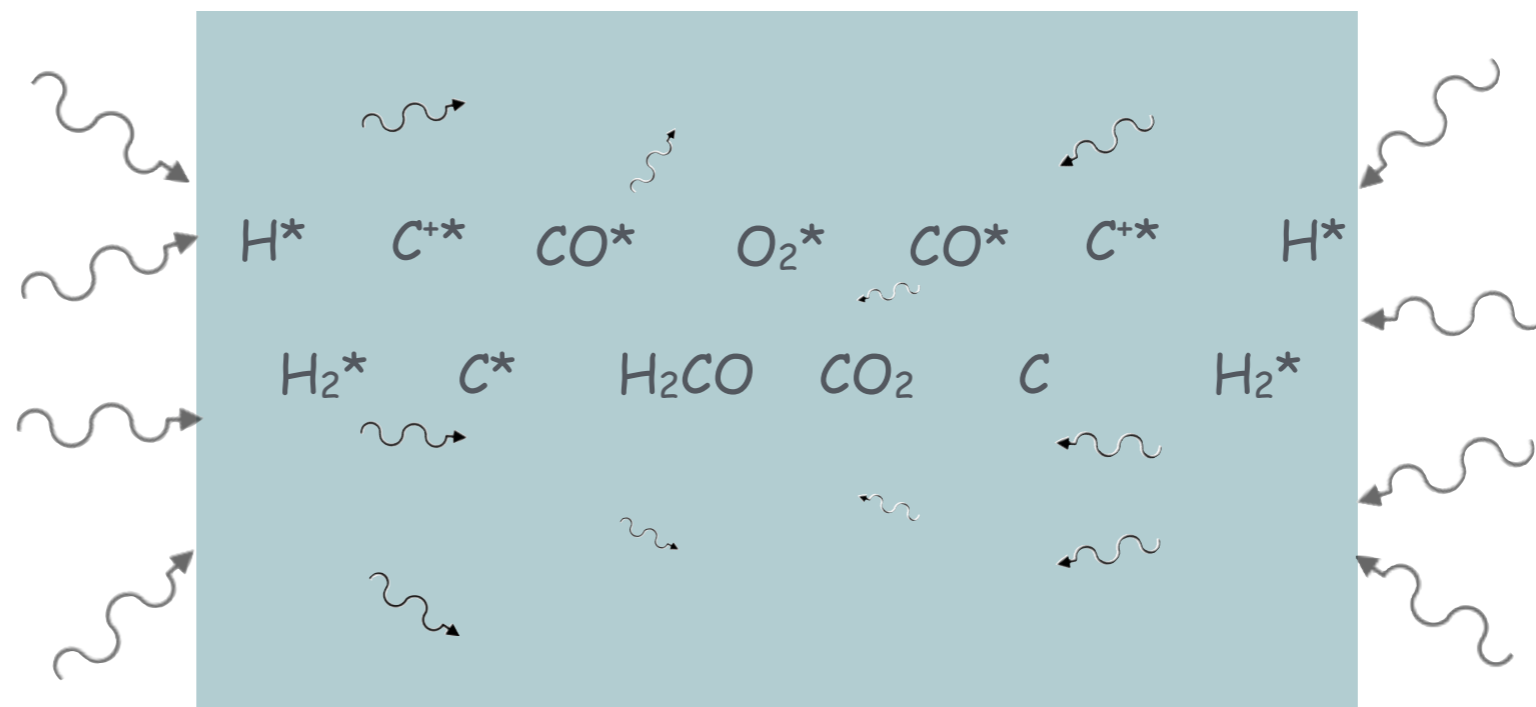


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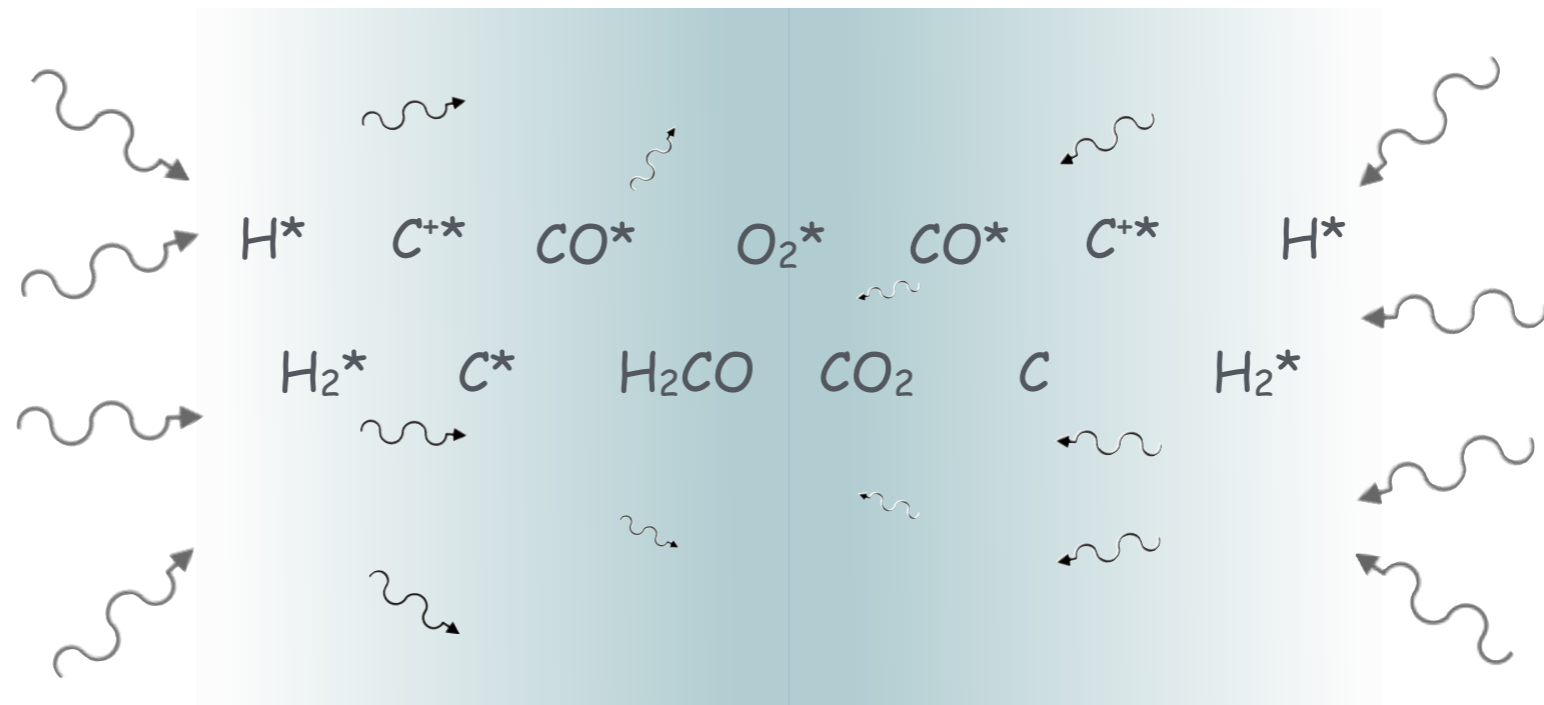


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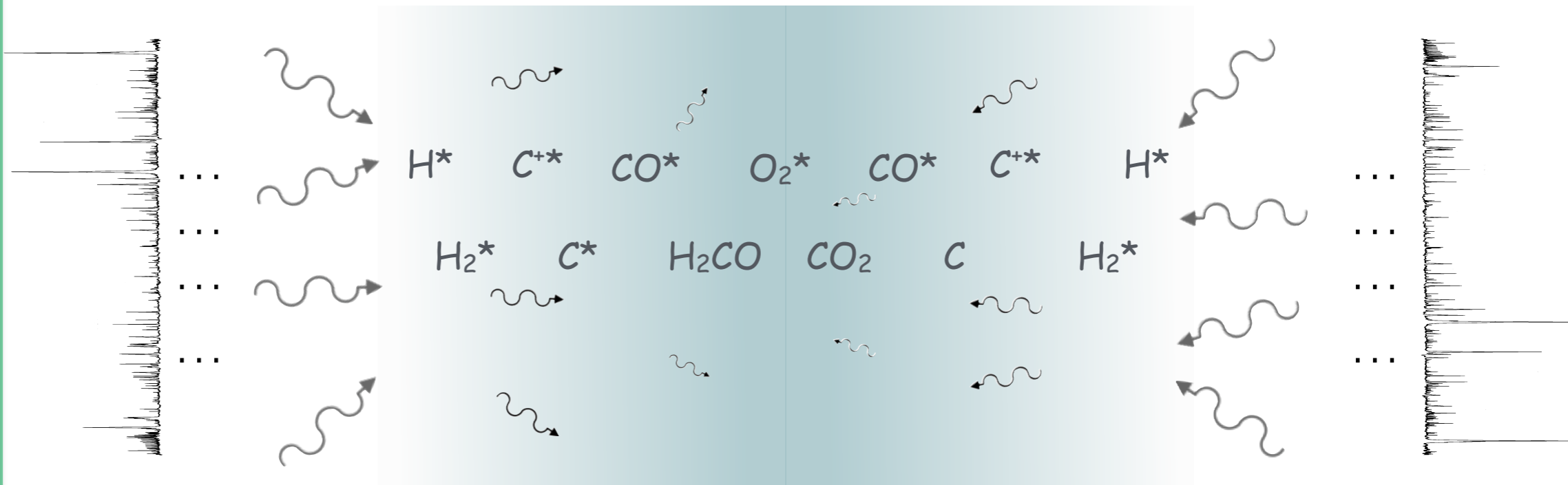


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- thermal balance
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- charge, emissivities
- excitation (thousands levels)
- tens thousands lines



## Limitations : 1D static

- chemistry  $\tau_{chem}^{i-n} \sim 3 \left( \frac{n_H}{10^3 \text{ cm}^{-3}} \right)^{-1} \text{ yr}$
- chemistry  $\tau_{chem}^{n-n} \sim 300 \left( \frac{n_H}{10^3 \text{ cm}^{-3}} \right)^{-1} \text{ yr}$
- ionization  $\tau_{CR} \sim 0.03 - 3 \text{ Gyr}$
- grains  $\tau_{gr}^{H_2} \sim 1 \left( \frac{n_H}{10^3 \text{ cm}^{-3}} \right)^{-1} \left( \frac{T}{50} \right)^{-1/2} \text{ Myr}$

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- 
- free fall  $\tau_{ff} \sim 0.14 \left( \frac{n_H}{10^3 \text{ cm}^{-3}} \right)^{-1/2} \text{ Myr}$
  - turnover  $\tau_{turn} \sim 3 \left( \frac{L}{3 \text{ pc}} \right) \left( \frac{u}{3 \text{ kms}^{-1}} \right)^{-1} \text{ Myr}$
  - turb dissipation  $\tau_{diss} \sim 10 \left( \frac{L}{3 \text{ pc}} \right)^{1/2} \left( \frac{u}{3 \text{ kms}^{-1}} \right)^{-3/2} \text{ yr}$
  - Ionization front  $t_{if} \sim 3 \times 10^5 \left( \frac{10^3}{n_H} \right) \left( \frac{N_0}{10^{21}} \right) \text{ yr}$

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## Limitations : 1D static

- transient processes not described
- loosing influence of initial conditions
  - ✓ loose understanding of physical evolution
  - ✓ bistability not treated
- no 3D structure
  - ✓ lack of realistic geometry
  - ✓ no permeability to UV photons
  - ✓ miss opacity effects

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## Advantages

- stress on microphysics
- asymptotic behavior
- avoid empirical laws  
⇒ results from first principles
- exact transfer  
⇒ non LTE, non LVG
- solve thousands couplings
- adaptative and versatile
- resolve chemical transitions  
 $H / H_2 \quad l_{H-H_2} \sim 100 \text{ AU}$   
 $C^+ / C \quad l_{C^+-C} \sim 0.5 \text{ pc}$
- use as virtual world
- large number of predictions
- exploration / grid of models
  - ✓ laws for simulations
  - ✓ efficient inversion tool

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model parameters  
radiation field, density, CR flux, elements, ...



initial conditions  
chemical & temperature profiles, ...

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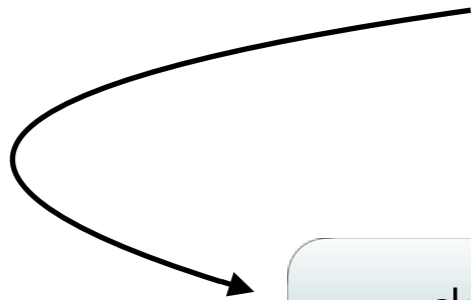
# Iterative procedure

model parameters  
radiation field, density, CR flux, elements, ...



initial conditions  
chemical & temperature profiles, ...

dust properties  
T, abs, emission



## Outline

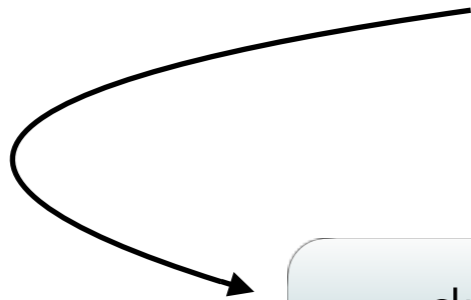
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dust properties  
T, abs, emission



radiative transfer  
interactions gas, dust

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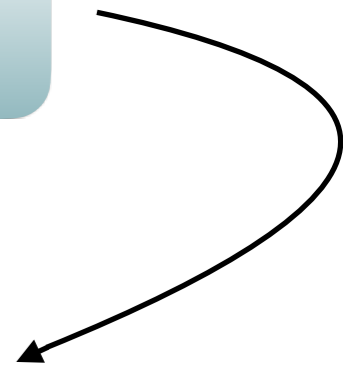
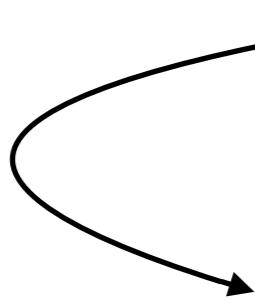
initial conditions  
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radiative transfer  
interactions gas, dust

dust charge



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initial conditions  
chemical & temperature profiles, ...

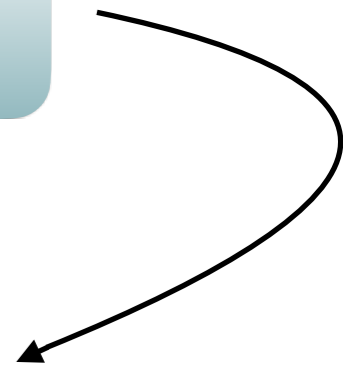
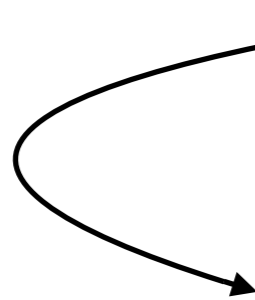
dust properties  
T, abs, emission



radiative transfer  
interactions gas, dust

dust charge

chemistry



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples



# Iterative procedure

model parameters  
radiation field, density, CR flux, elements, ...

initial conditions  
chemical & temperature profiles, ...

dust properties  
T, abs, emission

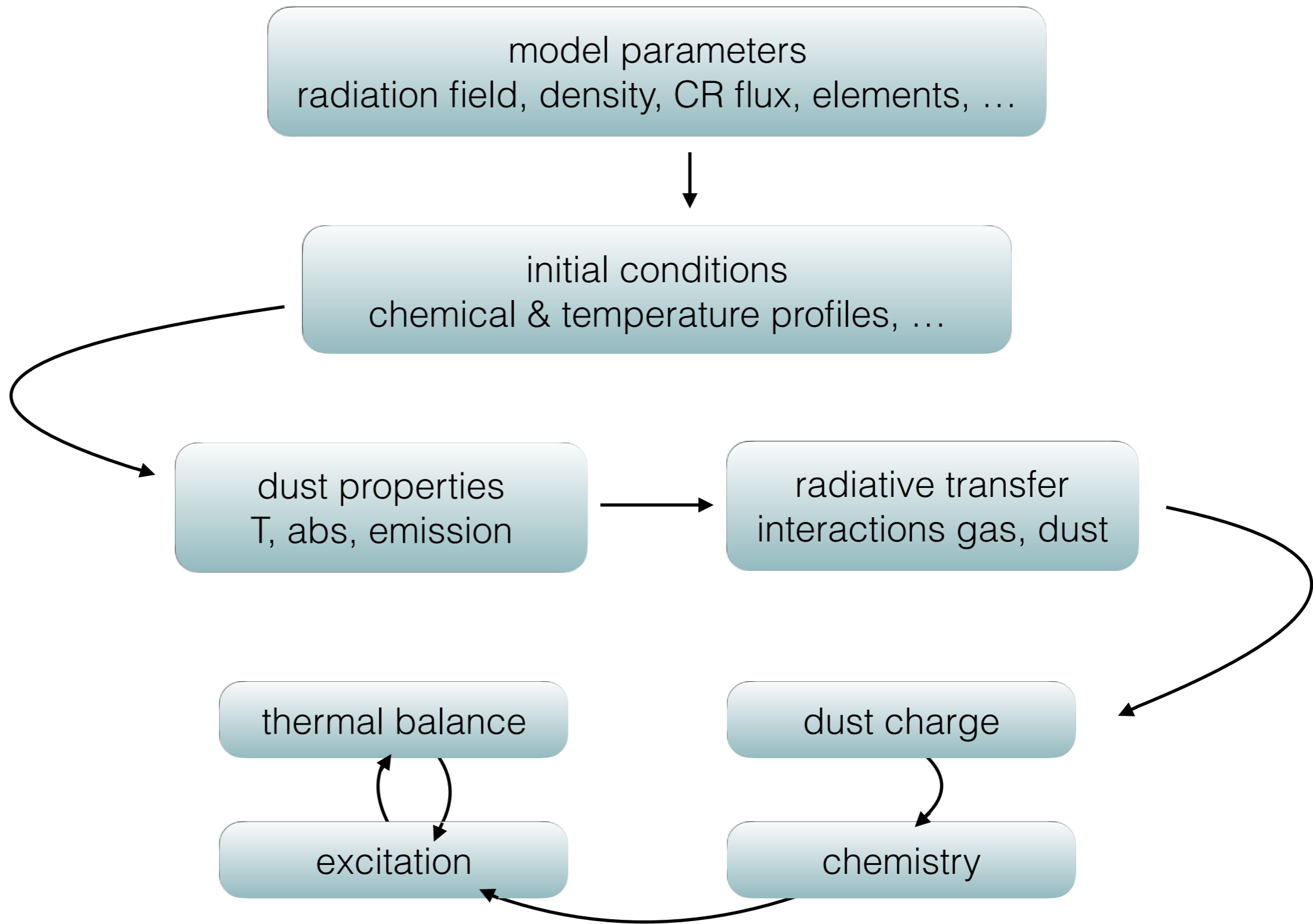
radiative transfer  
interactions gas, dust

thermal balance

dust charge

excitation

chemistry



## Outline

- introduction on PDRs
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- user guide first steps
- examples

# Iterative procedure

model parameters  
radiation field, density, CR flux, elements, ...

initial conditions  
chemical & temperature profiles, ...

dust properties  
T, abs, emission

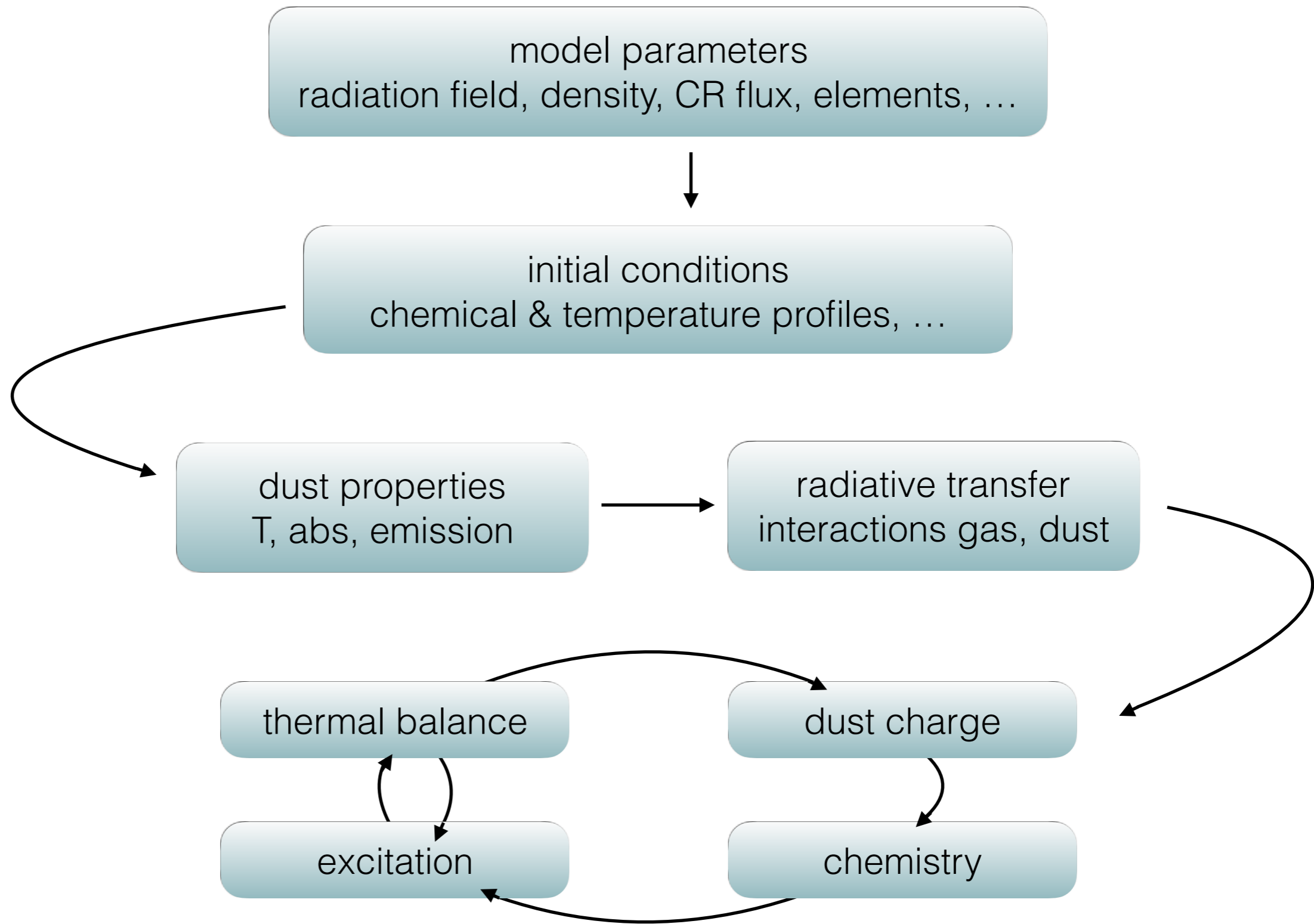
radiative transfer  
interactions gas, dust

thermal balance

dust charge

excitation

chemistry



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
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# Iterative procedure

model parameters  
radiation field, density, CR flux, elements, ...

initial conditions  
chemical & temperature profiles, ...

dust properties  
T, abs, emission

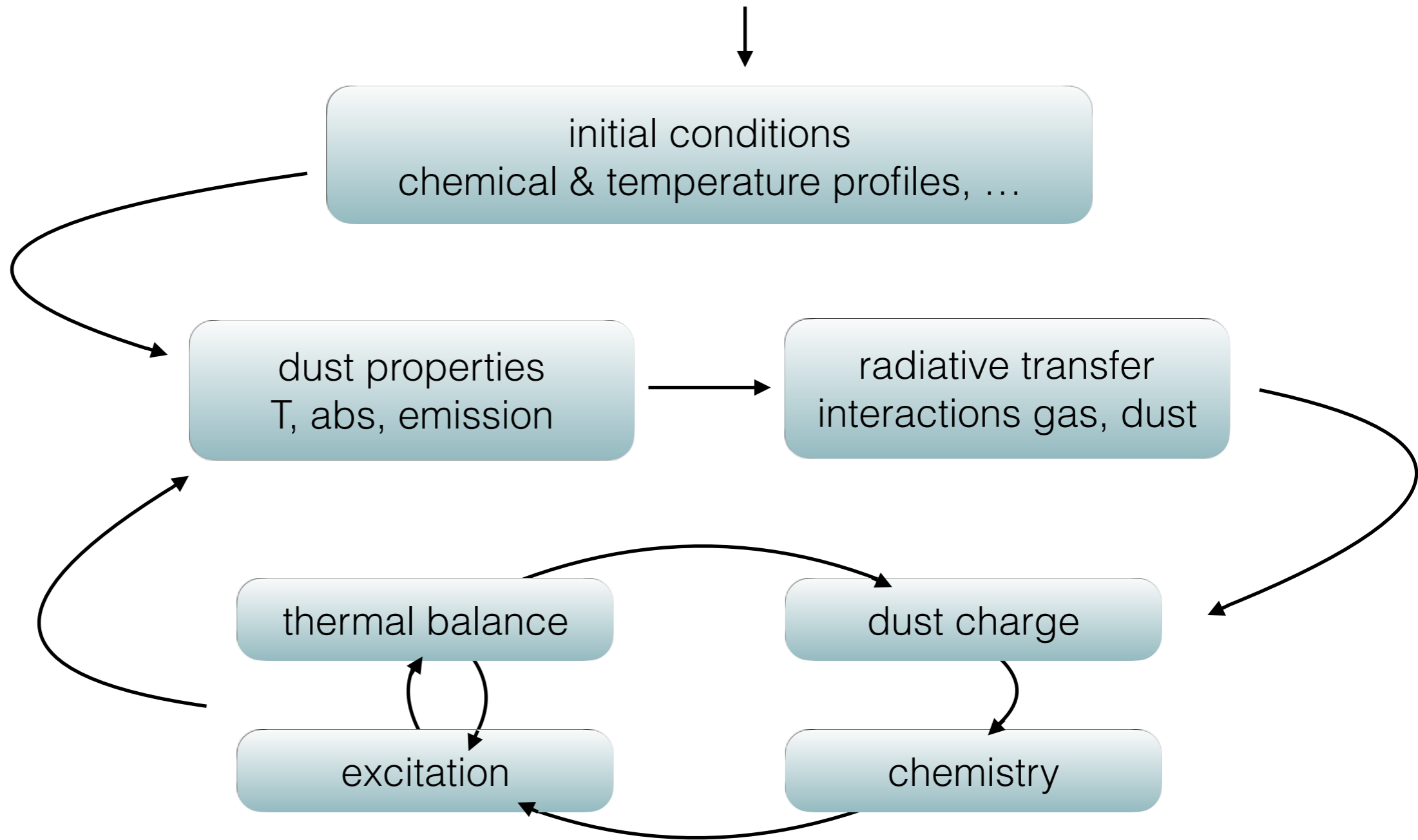
radiative transfer  
interactions gas, dust

thermal balance

dust charge

excitation

chemistry



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

## Computational time

⇒ Radiative transfer and statistical treatments are time consuming

⇒ Chemistry and thermal balance are fast

### Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

## Computational time

⇒ Radiative transfer and statistical treatments are time consuming

⇒ Chemistry and thermal balance are fast

version	Method	point in grid	time
PDRLight	dust absorption / emission	7 000	1/2 h
PDR 1.5.2	+ gas continuum absorption	13 000	4 h
PDR 1.5.2	+ H2 mutual lines shielding	27 000	9 h
PDR 1.5.2	+ CO mutual lines shielding	59 000	17 h
PDR 1.5.2	+ isotopes	77 000	26 h

## Outline

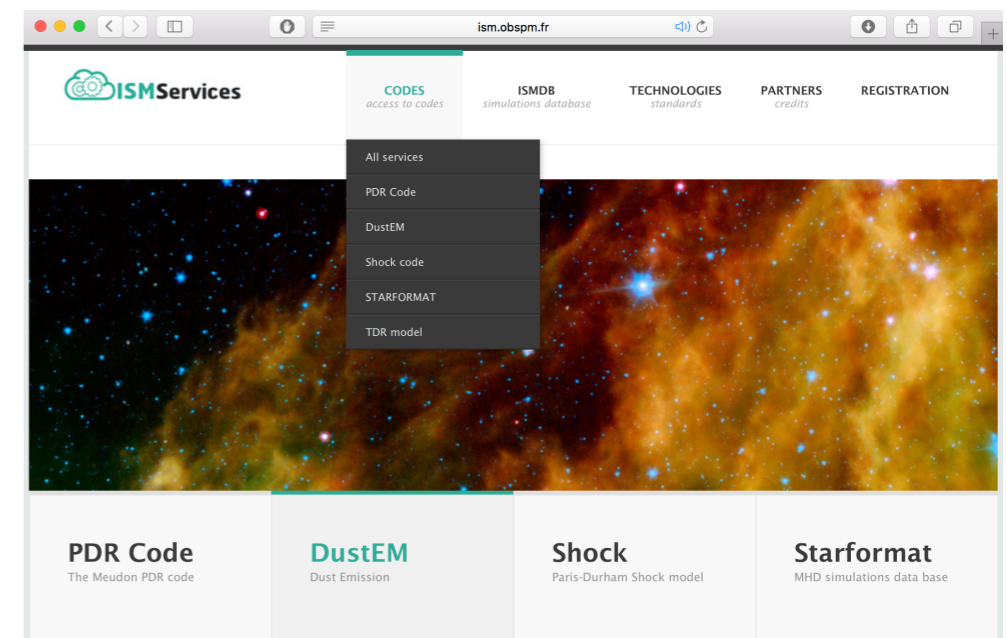
- introduction on PDRs
- assumptions
- user guide first steps
- examples



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

- download from the platform [ism.obspm.fr](http://ism.obspm.fr)  
*code & documentation*

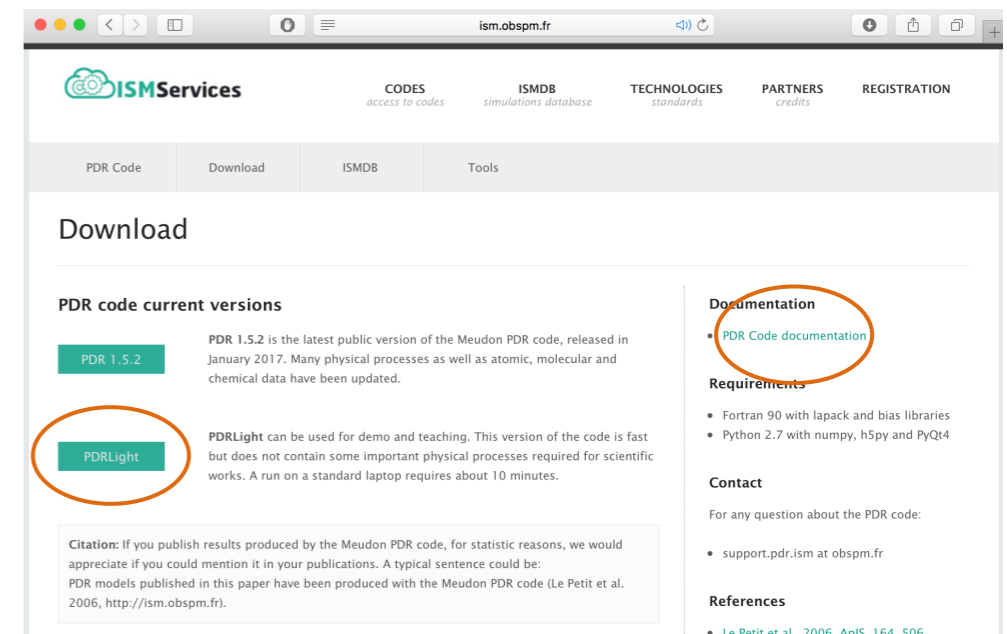
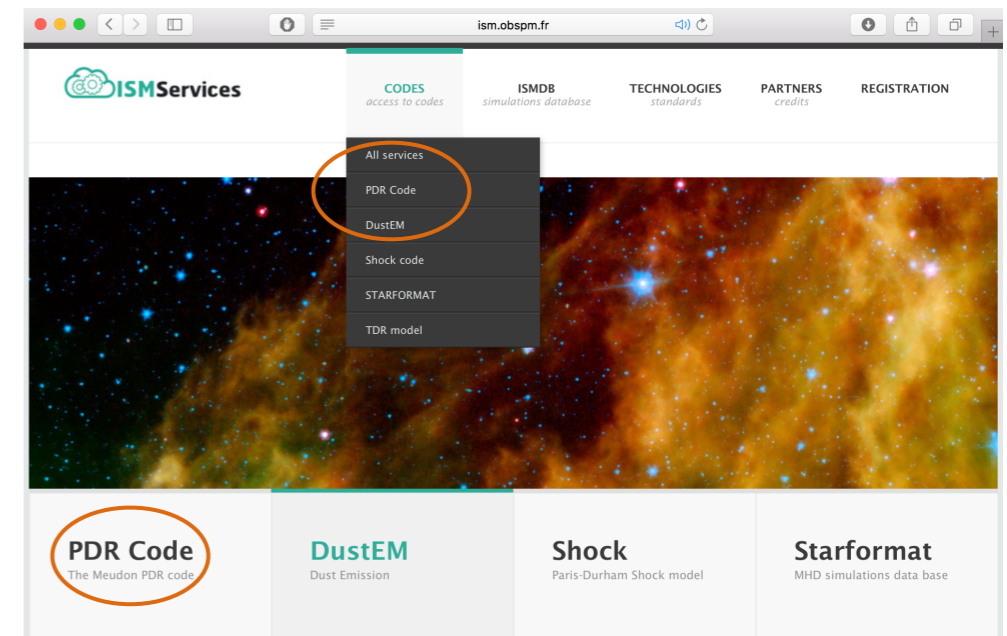


# ISM platform

- download from the platform  
[ism.obspm.fr](http://ism.obspm.fr)  
*code & documentation*

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

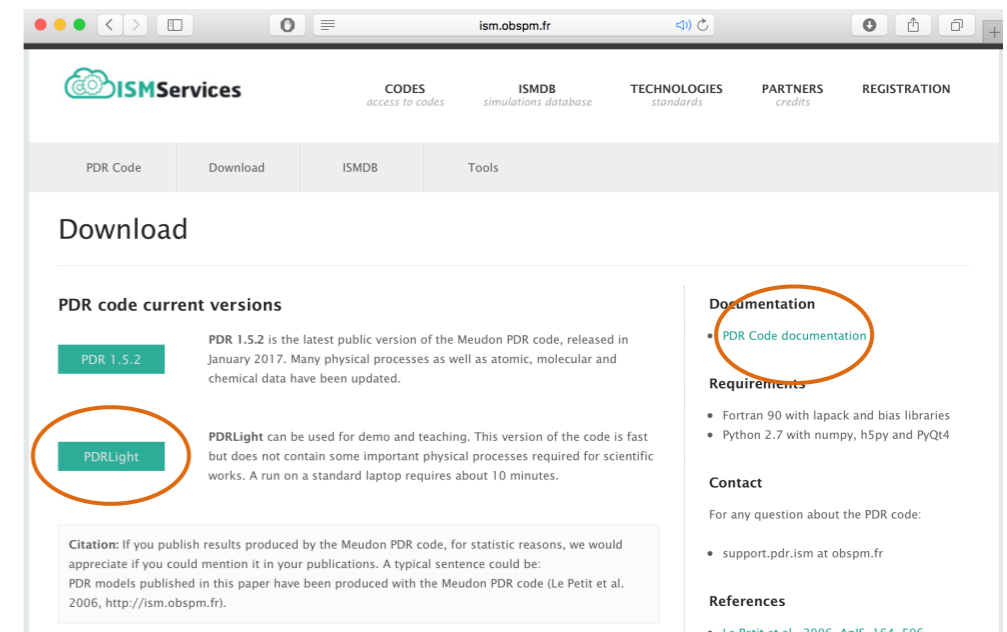
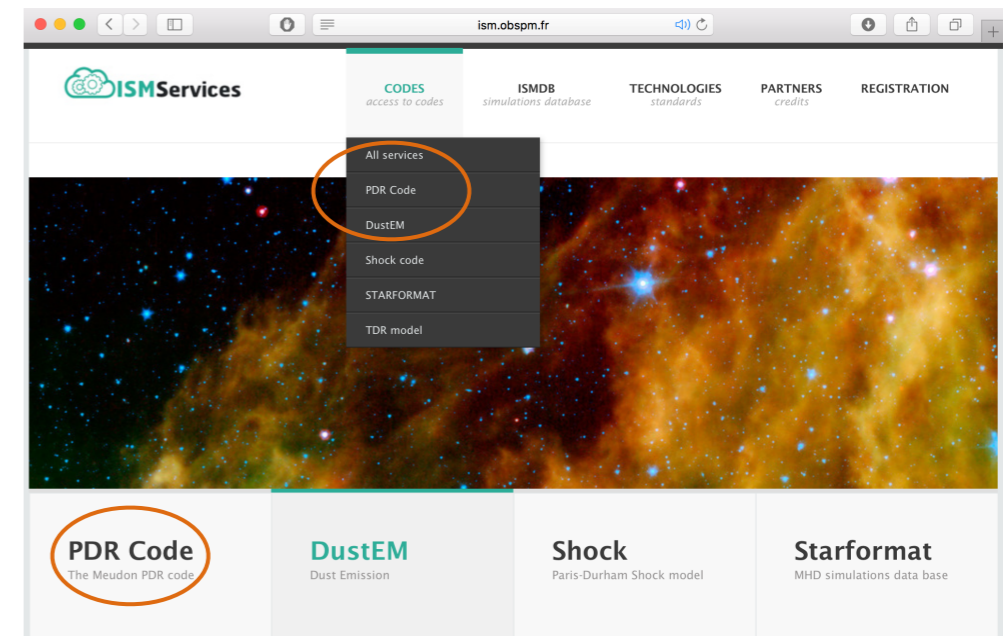


# ISM platform

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples








- download from the platform [ism.obspm.fr](http://ism.obspm.fr)  
*code & documentation*
- untar and compile  
*fortran compiler / librairies*  
*python2.7 / librairies*  
*web browser*
- prepare the input files  
*end criterion*  
*output options*
- run the model
- extract & analyse the results



# Inputs

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

	AnalysisTools	→	extraction / analysis tools
	data	→	input files
	out	→	output files
	PDRDoc.html	→	link to online doc
	PDRLight_1.0_svn	→	xcode project
	README	→	short help
	src	→	source code - compilation - run


# Inputs

 data

 Astrodata




astrophysical sources

 Chimie




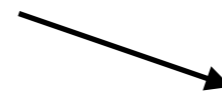
chemical networks

 Grains



grains abs / heat  
coefficients


 Collisions



 Levels



atom / mol data

 Lines



 pdr.in



main input file

 photodest.flag



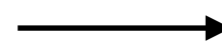
photodestr. param

 Sections



photodestr. cross sections

 spectre.flag



detailed balance param

 Techconfig



output configuration

 UVdata




























H / H<sub>2</sub> / CO UV lines

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples



# Inputs

 data		
 Astrodata		astrophysical sources
 Chimie		<b>chemical networks</b>
 Grains		grains abs / heat coefficients
 Collisions		
 Levels		atom / mol data
 Lines		
 pdr.in		<b>main input file</b>
 photodest.flag		photodestr. param
 Sections		photodestr. cross sections
 spectre.flag		<b>detailed balance param</b>
 Techconfig		output configuration
 UVdata		H / H <sub>2</sub> / CO UV lines

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

Main input file : `pdr.in`

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

```

~/Travail/Codes-Chimie/PDR_svn/1.5.2_light_svn/data/pdr.in
1   ExampleDiffuse      ! modele   : Output files radix-
2   ch1612_iso_Mathis.chi ! chimie   : Chemistry file (ex: ch
3   1                    ! ifafm    : Number of global itera
4   3.00e+00            ! Avmax    : Integration limit (Av)
5   1.00e+02            ! densh    : Initial density (nH =
6   1                    ! F_ISRF   : 1 = Mathis, 2 = Draine
7   1.00e+00            ! radm     : Radiation field intens
8   1.00e+00            ! radp     : Radiation field intens
9   none.txt            ! srcpp    : Additional plan parall
10  -0.000e+00          ! d_sour   : Star distance (pc) (<0
11  5.00e+00            ! fmrc     : Cosmic rays ionisation
12  1                    ! ieqth    : thermal Balance (1 : s
13  7.00e+01            ! tgaz     : Initial temperature (K
14  0                    ! ifisob   : State equation (0: nH
15  none.pfl            ! fprofil  : Temperature-density pr
16  1.00e+06            ! presse   : Initial Pressure (cm-3
17  2.00e+00            ! vturb    : turbulent velocity (Do
18  2                    ! ichh2    : H + H2 collision rate
19  Galaxy              ! los_ext  : Line of sight extincti
20  3.10                 ! rrr      : Rv = Av / E(B-V) (Typ
21  1.00e-00            ! metal    : metallicity Z (automat
22  5.80e+21            ! cdunit_0 : NH / E(B-V) for Z = 1,
23  1.00e-02            ! gratio_0 : Mass grains / mass gas
24  4.60e-02            ! q_pah    : PAH mass fraction (def
25  3.50e+00            ! alpgr    : grains distribution in
26  1.00e-07            ! rgrmin   : Grains minimum radius
27  3.00e-05            ! rgrmax   : Grains maximum radius
28  0                    ! F_DUST_P : 0 - Without DUSTEM, F&
29  0                    ! iforh2   : H2 formation on grains
30  4                    ! istic    : H2 sticking on grain m
31  0                    ! F_W_ALL_IFAF : 1 = Write output f
32

```

L: 32 C: 1 (none) Unicode (UTF-8) Unix (LF) 2 692 / 316 / 32 100%

Main input file : `pdr.in`

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

```
~/Travail/Codes-Chimie/PDR_svn/1.5.2_light_svn/data/pdr.in
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```

parameters descrip

Main input file : `pdr.in`

## Outline

- introduction on PDRs
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- user guide first steps
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```
~/Travail/Codes-Chimie/PDR_svn/1.5.2_light_svn/data/pdr.in
1  ExampleDiffuse      ! modele   : Output files radix-
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10 -0.000e+00         ! d_sour   : Star distance (pc) (<0
11 5.00e+00           ! fmrc     : Cosmic rays ionisation
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16 1.00e+06           ! presse   : Initial Pressure (cm-3
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parameters descrip

in & out files

Main input file : `pdr.in`

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7  1.00e+00           ! radm     : Radiation field intens
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9  none.txt           ! srcpp    : Additional plan parall
10 -0.000e+00         ! d_sour   : Star distance (pc) (<0
11 5.00e+00           ! fmrc     : Cosmic rays ionisation
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```

parameters descrip

in & out files

main parameters

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples



Main input file : `pdr.in`

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4  3.00e+00           ! Avmax    : Integration limit (Av)
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31 0                   ! F_W_ALL_IFAF : 1 = Write output f
32

```

parameters descrip

in & out files

main parameters

extinction curve

## Outline

- introduction on PDRs
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- examples

Main input file : `pdr.in`

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~/Travail/Codes-Chimie/PDR_svn/1.5.2_light_svn/data/pdr.in
1  ExampleDiffuse      ! modele  : Output files radix-
2  ch1612_iso_Mathis.chi ! chimie  : Chemistry file (ex: ch
3  1                   ! ifafm   : Number of global itera
4  3.00e+00           ! Avmax   : Integration limit (Av)
5  1.00e+02           ! densh   : Initial density (nH =
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7  1.00e+00           ! radm    : Radiation field intens
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9  none.txt           ! srcpp   : Additional plan parall
10 -0.000e+00         ! d_sour  : Star distance (pc) (<0
11 5.00e+00           ! fmrc    : Cosmic rays ionisation
12 1                   ! ieqth   : thermal Balance (1 : s
13 7.00e+01           ! tgaz    : Initial temperature (K
14 0                   ! ifisob  : State equation (0: nH
15 none.pfl           ! fprofil : Temperature-density pr
16 1.00e+06           ! presse  : Initial Pressure (cm-3
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18 2                   ! ichh2   : H + H2 collision rate
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21 1.00e-00           ! metal   : metallicity Z (automat
22 5.80e+21           ! cdunit_0 : NH / E(B-V) for Z = 1,
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25 3.50e+00           ! alpgr   : grains distribution in
26 1.00e-07           ! rgrmin  : Grains minimum radius
27 3.00e-05           ! rgrmax  : Grains maximum radius
28 0                   ! F_DUST_P : 0 - Without DUSTEM, F&
29 0                   ! iforh2  : H2 formation on grains
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32

```

parameters descrip

in & out files

main parameters

extinction curve

dust model

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

## Chemistry file

ch1612\_iso\_Mathis.chi

- sets species
  - ✓ composition
  - ✓ initial abundances
- sets reactions
  - ✓ reactants / products
  - ✓ type & rate
- consistency
  - ✓ for all species at least
    - 1 formation
    - 1 destruction
  - ✓ set elem abundance

```

(New Document)
1 #INCHI num espece hcnohe isot metos metos - abinit-rel enthalpf INCHIKeyis
2 #-----1-----2-----3-----4-----5-----6-----7-----
3 1 h 1 0000 0000 00000 000000 0 .800E+00 51.634 YZCKVEUIGOORGS-UH
4 2 d 0 0000 1000 00000 000000 0 .000E+00 52.520 NOINCHIFOUND
5 3 h2 2 0000 0000 00000 000000 0 .100E+00 .000 UFHFLCQGNINYRP-UH
6 4 hd 1 0000 1000 00000 000000 0 1.500E-05 0.079 NOINCHIFOUND
7 5 he 0 0001 0000 00000 000000 0 .100E+00 .000 SWQJXJJOGLNCZEY-UH
8 6 c 0 1000 0000 00000 000000 0 .000E+00 169.978 OKTJSMVPCPJKN-UH
9 7 c* 0 0000 0100 00000 000000 0 .000E+00 169.978 NOINCHIFOUND
10 8 ch 1 1000 0000 00000 000000 0 .000E+00 141.177 VRLIPUYDFBXWCH-UH
11
343 -
344 - - cosmic-ray direct processes-
345 -
346 jan11 h2 crp h h 1.00E-01 .00 .00 1
347 jan11 h crp h+ electr 4.60E-01 .00 .00 1
348 jan11 h2 crp h+ h electr 4.00E-02 .00 .00 1
349 jan11 h2 crp h2+ electr 9.60E-01 .00 .00 1
350
535 -
536 - - neutral-neutral reactions-
537 -
538 jan11 h n2 nh n 8.63E-11 .50 71457.00 4
539 jan11 h nh n h2 8.19E-11 .68 950.00 4
540 jan11 h no n oh 3.60E-10 .00 24912.00 4
541
2958 -
2959 - - photoprocesses (ISRF of Mathis et al 1983)-
2960 -
2961 jan11 ch photon c h 6.60E-10 .00 1.15 5
2962 jan11 ch2 photon ch h 4.90E-10 .00 1.67 5
2963 jan11 ch3 photon ch2 h 3.20E-10 .00 1.88 5
2964 jan11 ch3 photon ch h2 3.20E-10 .00 1.88 5
2965
3194 -
3195 - - Ions and electrons / Grains-
3196 -
3197 NEUTR h+ grain h 1.00E+00 0.00 0.00 14
3198 NEUTR he+ grain he 1.00E+00 0.00 0.00 14
3199 NEUTR c+ grain c 1.00E+00 0.00 0.00 14
3200 NEUTR o+ grain o 1.00E+00 0.00 0.00 14
3201 NEUTR n+ grain n 1.00E+00 0.00 0.00 14
3202 NEUTR s+ grain s 1.00E+00 0.00 0.00 14
3203 NEUTR si+ grain si 1.00E+00 0.00 0.00 14
    
```

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

## Chemistry file

ch1612\_iso\_Mathis.chi

- sets species
  - ✓ **composition**
  - ✓ initial abundances
- sets reactions
  - ✓ reactants / products
  - ✓ type & rate
- consistency
  - ✓ for all species at least
    - 1 formation
    - 1 destruction
  - ✓ set elem abundance

```

(New Document)
1 #INCHI num espcie hcnobe isot metos metos - abinit-rel enthalpf INCHIKeyis
2 #-----1-----2-----3-----4-----5-----6-----7-----
3 1 h 1 0000 0000 00000 000000 0 .800E+00 51.634 YZCKVEUIGOORGS-UH
4 2 d 0 0000 1000 00000 000000 0 .000E+00 52.520 NOINCHIFOUND
5 3 h2 2 0000 0000 00000 000000 0 .100E+00 .000 UFHFLCQGNINYRNP-UH
6 4 hd 1 0000 1000 00000 000000 0 1.500E-05 0.079 NOINCHIFOUND
7 5 he 0 0001 0000 00000 000000 0 .100E+00 .000 SWQJXJJOGLNCZEY-UH
8 6 c 0 1000 0000 00000 000000 0 .000E+00 169.978 OKTJSMVPCPJKN-UH
9 7 c* 0 0000 0100 00000 000000 0 .000E+00 169.978 NOINCHIFOUND
10 8 ch 1 1000 0000 00000 000000 0 .000E+00 141.177 VRLIPUYDFBXWCH-UH
11
343 -
344 - - cosmic-ray direct processes-
345 -
346 jan11 h2 crp h h 1.00E-01 .00 .00 1
347 jan11 h crp h+ electr 4.60E-01 .00 .00 1
348 jan11 h2 crp h+ h electr 4.00E-02 .00 .00 1
349 jan11 h2 crp h2+ electr 9.60E-01 .00 .00 1
350
535 -
536 - - neutral-neutral reactions-
537 -
538 jan11 h n2 nh n 8.63E-11 .50 71457.00 4
539 jan11 h nh n h2 8.19E-11 .68 950.00 4
540 jan11 h no n oh 3.60E-10 .00 24912.00 4
541
2958 -
2959 - - photoprocesses (ISRF of Mathis et al 1983)-
2960 -
2961 jan11 ch photon c h 6.60E-10 .00 1.15 5
2962 jan11 ch2 photon ch h 4.90E-10 .00 1.67 5
2963 jan11 ch3 photon ch2 h 3.20E-10 .00 1.88 5
2964 jan11 ch3 photon ch h2 3.20E-10 .00 1.88 5
2965
3194 -
3195 - - Ions and electrons / Grains-
3196 -
3197 NEUTR h+ grain h 1.00E+00 0.00 0.00 14
3198 NEUTR he+ grain he 1.00E+00 0.00 0.00 14
3199 NEUTR c+ grain c 1.00E+00 0.00 0.00 14
3200 NEUTR o+ grain o 1.00E+00 0.00 0.00 14
3201 NEUTR n+ grain n 1.00E+00 0.00 0.00 14
3202 NEUTR s+ grain s 1.00E+00 0.00 0.00 14
3203 NEUTR si+ grain si 1.00E+00 0.00 0.00 14
    
```

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples



## Chemistry file

ch1612\_iso\_Mathis.chi

- sets species
  - ✓ composition
  - ✓ **initial abundances**
- sets reactions
  - ✓ reactants / products
  - ✓ type & rate
- consistency
  - ✓ for all species at least
    - 1 formation
    - 1 destruction
  - ✓ set elem abundance

(New Document)

1	#INCHI	num	espece	hcnohe	isot	metos	metos	abinit-rel	enthalpf	INCHIKeyis
2	#	1	2	3	4	5	6	7		
3	1	h	1	0000	0000	000000	0	.800E+00	51.634	YZCKVEUIGOORGS-UH
4	2	d	0	0000	1000	000000	0	.000E+00	52.520	NOINCHIFOUND
5	3	h2	2	0000	0000	000000	0	.100E+00	.000	UFHFLCQGNINYRP-UH
6	4	hd	1	0000	1000	000000	0	1.500E-05	0.079	NOINCHIFOUND
7	5	he	0	0001	0000	000000	0	.100E+00	.000	SWQJXJJOGLNCZEY-UH
8	6	c	0	1000	0000	000000	0	.000E+00	169.978	OKTJSMVPCPJKN-UH
9	7	c*	0	0000	0100	000000	0	.000E+00	169.978	NOINCHIFOUND
10	8	ch	1	1000	0000	000000	0	.000E+00	141.177	VRLIPUYDFBXWCH-UH
11	---									
343	-									
344	- cosmic-ray direct processes-									
345	---									
346	jan11	h2	crp	h	h			1.00E-01	.00	.00 1
347	jan11	h	crp	h+	electr			4.60E-01	.00	.00 1
348	jan11	h2	crp	h+	h	electr		4.00E-02	.00	.00 1
349	jan11	h2	crp	h2+	electr			9.60E-01	.00	.00 1
350	---									
535	-									
536	- neutral-neutral reactions-									
537	---									
538	jan11	h	n2	nh	n			8.63E-11	.50	71457.00 4
539	jan11	h	nh	n	h2			8.19E-11	.68	950.00 4
540	jan11	h	no	n	oh			3.60E-10	.00	24912.00 4
541	---									
2958	-									
2959	- photoprocesses (ISRF of Mathis et al 1983)-									
2960	---									
2961	jan11	ch	photon	c	h			6.60E-10	.00	1.15 5
2962	jan11	ch2	photon	ch	h			4.90E-10	.00	1.67 5
2963	jan11	ch3	photon	ch2	h			3.20E-10	.00	1.88 5
2964	jan11	ch3	photon	ch	h2			3.20E-10	.00	1.88 5
2965	---									
3194	-									
3195	- Ions and electrons / Grains-									
3196	---									
3197	NEUTR	h+	grain	h				1.00E+00	0.00	0.00 14
3198	NEUTR	he+	grain	he				1.00E+00	0.00	0.00 14
3199	NEUTR	c+	grain	c				1.00E+00	0.00	0.00 14
3200	NEUTR	o+	grain	o				1.00E+00	0.00	0.00 14
3201	NEUTR	n+	grain	n				1.00E+00	0.00	0.00 14
3202	NEUTR	s+	grain	s				1.00E+00	0.00	0.00 14
3203	NEUTR	si+	grain	si				1.00E+00	0.00	0.00 14

L: 3204 C: 1 (none) Unicode (UTF-8) Unix (LF) (never saved) 83 / 9 / 1 100%

## Outline

- introduction on PDRs
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## Chemistry file

ch1612\_iso\_Mathis.chi

- sets species
  - ✓ composition
  - ✓ initial abundances
- sets reactions
  - ✓ **reactants / products**
  - ✓ type & rate
- consistency
  - ✓ for all species at least
    - 1 formation
    - 1 destruction
  - ✓ set elem abundance

```

(New Document)
1 #INCHI num espcie hcnohe isot metos metos - abinit-rel enthalpf INCHIKeyis
2 #-----1-----2-----3-----4-----5-----6-----7-----
3 1 h 1 0000 0000 00000 000000 0 .800E+00 51.634 YZCKVEUIGOORGS-UH
4 2 d 0 0000 1000 00000 000000 0 .000E+00 52.520 NOINCHIFOUND
5 3 h2 2 0000 0000 00000 000000 0 .100E+00 .000 UFHFLCQGNINYRNP-UH
6 4 hd 1 0000 1000 00000 000000 0 1.500E-05 0.079 NOINCHIFOUND
7 5 he 0 0001 0000 00000 000000 0 .100E+00 .000 SWQJXJJOGLNCZEY-UH
8 6 c 0 1000 0000 00000 000000 0 .000E+00 169.978 OKTJSMVPCPJKN-UH
9 7 c* 0 0000 0100 00000 000000 0 .000E+00 169.978 NOINCHIFOUND
10 8 ch 1 1000 0000 00000 000000 0 .000E+00 141.177 VRLIPUYDFBXWCH-UH
11
343 -
344 - - cosmic-ray direct processes-
345 -
346 jan11 h2 crp h h 1.00E-01 .00 .00 1
347 jan11 h crp h+ electr 4.60E-01 .00 .00 1
348 jan11 h2 crp h+ h electr 4.00E-02 .00 .00 1
349 jan11 h2 crp h2+ electr 9.60E-01 .00 .00 1
350
535 -
536 - - neutral-neutral reactions-
537 -
538 jan11 h n2 nh n 8.63E-11 .50 71457.00 4
539 jan11 h nh n h2 8.19E-11 .68 950.00 4
540 jan11 h no n oh 3.60E-10 .00 24912.00 4
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2958 -
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2960 -
2961 jan11 ch photon c h 6.60E-10 .00 1.15 5
2962 jan11 ch2 photon ch h 4.90E-10 .00 1.67 5
2963 jan11 ch3 photon ch2 h 3.20E-10 .00 1.88 5
2964 jan11 ch3 photon ch h2 3.20E-10 .00 1.88 5
2965
3194 -
3195 - - Ions and electrons / Grains-
3196 -
3197 NEUTR h+ grain h 1.00E+00 0.00 0.00 14
3198 NEUTR he+ grain he 1.00E+00 0.00 0.00 14
3199 NEUTR c+ grain c 1.00E+00 0.00 0.00 14
3200 NEUTR o+ grain o 1.00E+00 0.00 0.00 14
3201 NEUTR n+ grain n 1.00E+00 0.00 0.00 14
3202 NEUTR s+ grain s 1.00E+00 0.00 0.00 14
3203 NEUTR si+ grain si 1.00E+00 0.00 0.00 14
    
```

## Outline

- introduction on PDRs
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## Chemistry file

ch1612\_iso\_Mathis.chi

- sets species
  - ✓ composition
  - ✓ initial abundances
- sets reactions
  - ✓ reactants / products
  - ✓ **type & rate**
- consistency
  - ✓ for all species at least
    - 1 formation
    - 1 destruction
  - ✓ set elem abundance

```
(New Document)
1 #INCHI num espece hcnohe isot metos metos - abinit-rel enthalpf INCHIKeyis
2 #-----1-----2-----3-----4-----5-----6-----7-----
3 1 h 1 0000 0000 00000 000000 0 .800E+00 51.634 YZCKVEUIGOORGS-UH
4 2 d 0 0000 1000 00000 000000 0 .000E+00 52.520 NOINCHIFOUND
5 3 h2 2 0000 0000 00000 000000 0 .100E+00 .000 UFHFLCQGNINYRNP-UH
6 4 hd 1 0000 1000 00000 000000 0 1.500E-05 0.079 NOINCHIFOUND
7 5 he 0 0001 0000 00000 000000 0 .100E+00 .000 SWQJXJQGLNCZEY-UH
8 6 c 0 1000 0000 00000 000000 0 .000E+00 169.978 OKTJSMVPCPJKN-UH
9 7 c* 0 0000 0100 00000 000000 0 .000E+00 169.978 NOINCHIFOUND
10 8 ch 1 1000 0000 00000 000000 0 .000E+00 141.177 VRLIPUYDFBXWCH-UH
11
343 -
344 - - cosmic-ray direct processes-
345 -
346 jan11 h2 crp h h 1.00E-01 .00 .00 1
347 jan11 h crp h+ electr 4.60E-01 .00 .00 1
348 jan11 h2 crp h+ h electr 4.00E-02 .00 .00 1
349 jan11 h2 crp h2+ electr 9.60E-01 .00 .00 1
350
535 -
536 - - neutral-neutral reactions-
537 -
538 jan11 h n2 nh n 8.63E-11 .50 71457.00 4
539 jan11 h nh n h2 8.19E-11 .68 950.00 4
540 jan11 h no n oh 3.60E-10 .00 24912.00 4
541
2958 -
2959 - - photoprocesses (ISRF of Mathis et al 1983)-
2960 -
2961 jan11 ch photon c h 6.60E-10 .00 1.15 5
2962 jan11 ch2 photon ch h 4.90E-10 .00 1.67 5
2963 jan11 ch3 photon ch2 h 3.20E-10 .00 1.88 5
2964 jan11 ch3 photon ch h2 3.20E-10 .00 1.88 5
2965
3194 -
3195 - - Ions and electrons / Grains-
3196 -
3197 NEUTR h+ grain h 1.00E+00 0.00 0.00 14
3198 NEUTR he+ grain he 1.00E+00 0.00 0.00 14
3199 NEUTR c+ grain c 1.00E+00 0.00 0.00 14
3200 NEUTR o+ grain o 1.00E+00 0.00 0.00 14
3201 NEUTR n+ grain n 1.00E+00 0.00 0.00 14
3202 NEUTR s+ grain s 1.00E+00 0.00 0.00 14
3203 NEUTR si+ grain si 1.00E+00 0.00 0.00 14
L: 3204 C: 1 (none) Unicode (UTF-8) Unix (LF) (never saved) 83 / 9 / 1 100%
```

spectre.flag

```
~/Travail/Codes-Chimie/P.../.../data/spectre.flag
1  #Activation(1: yes,0:no) Species nused nus
2  1 h          -1   1  1.0 !-
3  1 h2        250  50  2.0 !-
4  1 hd         9   9  3.0 !-
5  1 co         64  10 28.0 ! 12C 160-
6  1 c*o        32  10 29.0 ! 13C 160-
7  1 co*        32  10 30.0 ! 12C 180-
8  1 c*o*       31  10 31.0 ! 13C 180-
9  1 c          -1   3 12.0 !-
10 1 n          -1   1 14.0 !-
11 1 o          -1   3 16.0 ! Set 5,5 to conside
12 1 o2         -1  24 32.0 !-
13 1 s          -1   3 32.0 !-
14 1 si         -1   3 28.0 !-
15 1 cs         -1  10 44.0 !-
16 1 hcn        -1  10 27.0 !-
17 1 oh         -1  20 17.0 !-
18 1 h2o        -1  15 18.0 !-
19 1 c+         -1   2 12.0 !-
20 0 c++        -1   1 12.0 !-
21 0 n+         -1   1 14.0 !-
22 0 n++        -1   2 14.0 !-
23 0 o+         -1   1 16.0 !-
24 0 o++        -1   3 16.0 !-
25 0 s+         -1   1 32.0 !-
26 0 s++        -1   3 32.0 !-
27 1 si+        -1   2 28.0 !-

L: 1 C: 1 (none) Unicode (UTF-8) Unix (LF)
```

Detailed balance parameters



- thermal balance
- atom / mol level excitation
- adaptative method

## Outline

- introduction on PDRs
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spectre.flag

Detailed balance parameters

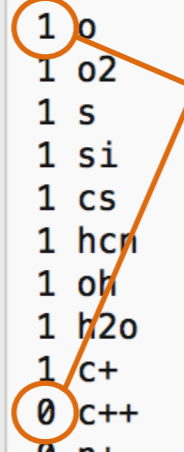


- thermal balance
- atom / mol level excitation
- adaptative method

use or drop species

```

~/Travail/Codes-Chimie/P.../.../data/spectre.flag
1  #Activation(1: yes,0:no) Species nused nus
2  1 h      -1  1  1.0 !-
3  1 h2     250 50  2.0 !-
4  1 hd      9  9  3.0 !-
5  1 co      64 10 28.0 ! 12C 160-
6  1 c*o     32 10 29.0 ! 13C 160-
7  1 co*     32 10 30.0 ! 12C 180-
8  1 c*o*    31 10 31.0 ! 13C 180-
9  1 c       -1  3 12.0 !-
10 1 n       -1  1 14.0 !-
11 1 o       -1  3 16.0 ! Set 5,5 to consid
12 1 o2      -1 24 32.0 !-
13 1 s       -1  3 32.0 !-
14 1 si      -1  3 28.0 !-
15 1 cs      -1 10 44.0 !-
16 1 hcn     -1 10 27.0 !-
17 1 oh      -1 20 17.0 !-
18 1 h2o     -1 15 18.0 !-
19 1 c+      -1  2 12.0 !-
20 0 c++     -1  1 12.0 !-
21 0 n+      -1  1 14.0 !-
22 0 n++     -1  2 14.0 !-
23 0 o+      -1  1 16.0 !-
24 0 o++     -1  3 16.0 !-
25 0 s+      -1  1 32.0 !-
26 0 s++     -1  3 32.0 !-
27 1 si+     -1  2 28.0 !-
    
```



## Outline

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spectre.flag

Detailed balance parameters



- thermal balance
- atom / mol level excitation
- adaptative method

use or drop species

use at least 10 lev of CO

use at most 64 lev of CO

```

~/Travail/Codes-Chimie/P.../.../data/spectre.flag
1  #Activation(1: yes,0:no) Species nused nus
2  1 h      -1  1  1.0 !-
3  1 h2     250 50  2.0 !-
4  1 hd      9  9  3.0 !-
5  1 co     64 10 28.0 ! 12C 160-
6  1 c*o    32 10 29.0 ! 13C 160-
7  1 co*    32 10 30.0 ! 12C 180-
8  1 c*o*   31 10 31.0 ! 13C 180-
9  1 c      -1  3 12.0 !-
10 1 n      -1  1 14.0 !-
11 1 o      -1  3 16.0 ! Set 5,5 to consid
12 1 o2     -1 24 32.0 !-
13 1 s      -1  3 32.0 !-
14 1 si     -1  3 28.0 !-
15 1 cs     -1 10 44.0 !-
16 1 hcn    -1 10 27.0 !-
17 1 oh     -1 20 17.0 !-
18 1 h2o    -1 15 18.0 !-
19 1 c+     -1  2 12.0 !-
20 0 c++    -1  1 12.0 !-
21 0 n+     -1  1 14.0 !-
22 0 n++    -1  2 14.0 !-
23 0 o+     -1  1 16.0 !-
24 0 o++    -1  3 16.0 !-
25 0 s+     -1  1 32.0 !-
26 0 s++    -1  3 32.0 !-
27 1 si+    -1  2 28.0 !-
    
```

L: 1 C: 1 (none) Unicode (UTF-8) Unix (LF)

## Outline

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spectre.flag

Detailed balance parameters



- thermal balance
- atom / mol level excitation
- adaptative method

use or drop species

use at least 10 lev of CO

use at most 64 lev of CO

use all levels of C+

```

~/Travail/Codes-Chimie/P.../.../data/spectre.flag
1  #Activation(1: yes,0:no) Species nused nus
2  1 h      -1  1  1.0 !-
3  1 h2     250 50  2.0 !-
4  1 hd      9  9  3.0 !-
5  1 co     64 10 28.0 ! 12C 160-
6  1 c*o    32 10 29.0 ! 13C 160-
7  1 co*    32 10 30.0 ! 12C 180-
8  1 c*o*   31 10 31.0 ! 13C 180-
9  1 c      -1  3 12.0 !-
10 1 n      -1  1 14.0 !-
11 1 o      -1  3 16.0 ! Set 5,5 to consid
12 1 o2     -1 24 32.0 !-
13 1 s      -1  3 32.0 !-
14 1 si     -1  3 28.0 !-
15 1 cs     -1 10 44.0 !-
16 1 hcn    -1 10 27.0 !-
17 1 oh     -1 20 17.0 !-
18 1 h2o    -1 15 18.0 !-
19 1 c+     -1  2 12.0 !-
20 0 c++    -1  1 12.0 !-
21 0 n+     -1  1 14.0 !-
22 0 n++    -1  2 14.0 !-
23 0 o+     -1  1 16.0 !-
24 0 o++    -1  3 16.0 !-
25 0 s+     -1  1 32.0 !-
26 0 s++    -1  3 32.0 !-
27 1 si+    -1  2 28.0 !-
    
```

## Outline

- introduction on PDRs
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- examples

## Makefile (in the `src/` directory)

- set the compiler name
- set compiler's options
- set libraries' paths
- compile

\$ make

→ PDR

```

43 #=====
44 #                                IFOPT (Intel)                                #
45 #=====
46 # FC = ifort-
47 -
48 #- Version 14.0-
49 # FFLAGS = -g -O2 -parallel -mkl -xSSSE3 -axSSSE3 -m64-
50 # FFLAGS = -g3 -O0 -mkl -check all -warn all -check:noarg_temp_created -traceback
51 -
52 # LIBS1 = -lpthread-
53 -
54 #=====
55 #                                GFORTTRAN                                #
56 #=====
57 FC = gfortran-
58 -
59 # Options and Path-
60 # FFLAGS = -g3 -O0 -fno-second-underscore -Wall -Wextra -Wno-compare-reals -fcheck
61 # FFLAGS = -g -Ofast -fno-second-underscore -falign-loops=16 -march=corei7 -ffast
62 # FFLAGS = -g -O2 -fno-second-underscore -falign-loops=16 -march=corei7 -ffast
63 FFLAGS = -g -O2 -fno-second-underscore -falign-loops=16 -ffast-math -fassoci
64 -
65 LIBS1 = -llapack -lblas-
66 -
67 #=====
68 #                                G95                                #
69 #=====
70 # FC = g95-
71 -
72 # Options and Path-
73 # FFLAGS = -O2 -fno-second-underscore-
74 # FFLAGS = -g3 -fno-second-underscore -Wall -Wextra -fbounds-check -ftrace=full -
75 # FFLAGS = -g3 -gdwarf-2 -fno-second-underscore -Wall -Wextra -fbounds-check -ftr
76 -
77 # LIBS1 = -llapack -lblas-
78 -

```

L: 1 C: 1    Make    Unicode (UTF-8)    Unix (LF)    Saved: 13/06/2017 10:55:28    8 374 / 665 / 166

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

Makefile (in the `src/` directory)

- set the compiler name
- set compiler's options
- set libraries' paths
- compile

```
$ make
```

```
→ PDR
```

```

43 #=====
44 #                                IFOPT (Intel)                                #
45 #=====
46 # FC = ifort-
47 -
48 #- Version 14.0-
49 # FFLAGS = -g -O2 -parallel -mkl -xSSSE3 -axSSSE3 -m64-
50 # FFLAGS = -g3 -O0 -mkl -check all -warn all -check:noarg_temp_created -tracebac
51 -
52 # LIBS1 = -lpthread-
53 -
54 #=====
55 #                                GFORTTRAN                                #
56 #=====
57 FC = gfortran-
58 -
59 # Options and Path-
60 # FFLAGS = -g3 -O0 -fno-second-underscore -Wall -Wextra -Wno-compare-reals -fchec
61 # FFLAGS = -g -Ofast -fno-second-underscore -falign-loops=16 -march=corei7 -ffast
62 # FFLAGS = -g -O2 -fno-second-underscore -falign-loops=16 -march=corei7 -ffast
63 FFLAGS = -g -O2 -fno-second-underscore -falign-loops=16 -ffast-math -fassoci
64 -
65 LIBS1 = -llapack -lblas-
66 -
67 #=====
68 #                                G95                                #
69 #=====
70 # FC = g95-
71 -
72 # Options and Path-
73 # FFLAGS = -O2 -fno-second-underscore-
74 # FFLAGS = -g3 -fno-second-underscore -Wall -Wextra -fbounds-check -ftrace=full -
75 # FFLAGS = -g3 -gdwarf-2 -fno-second-underscore -Wall -Wextra -fbounds-check -ftr
76 -
77 # LIBS1 = -llapack -lblas-
78 -

```

Run (in the `src/` directory)

- with default input file (`pdr.in`)

```
$ ./PDR
```

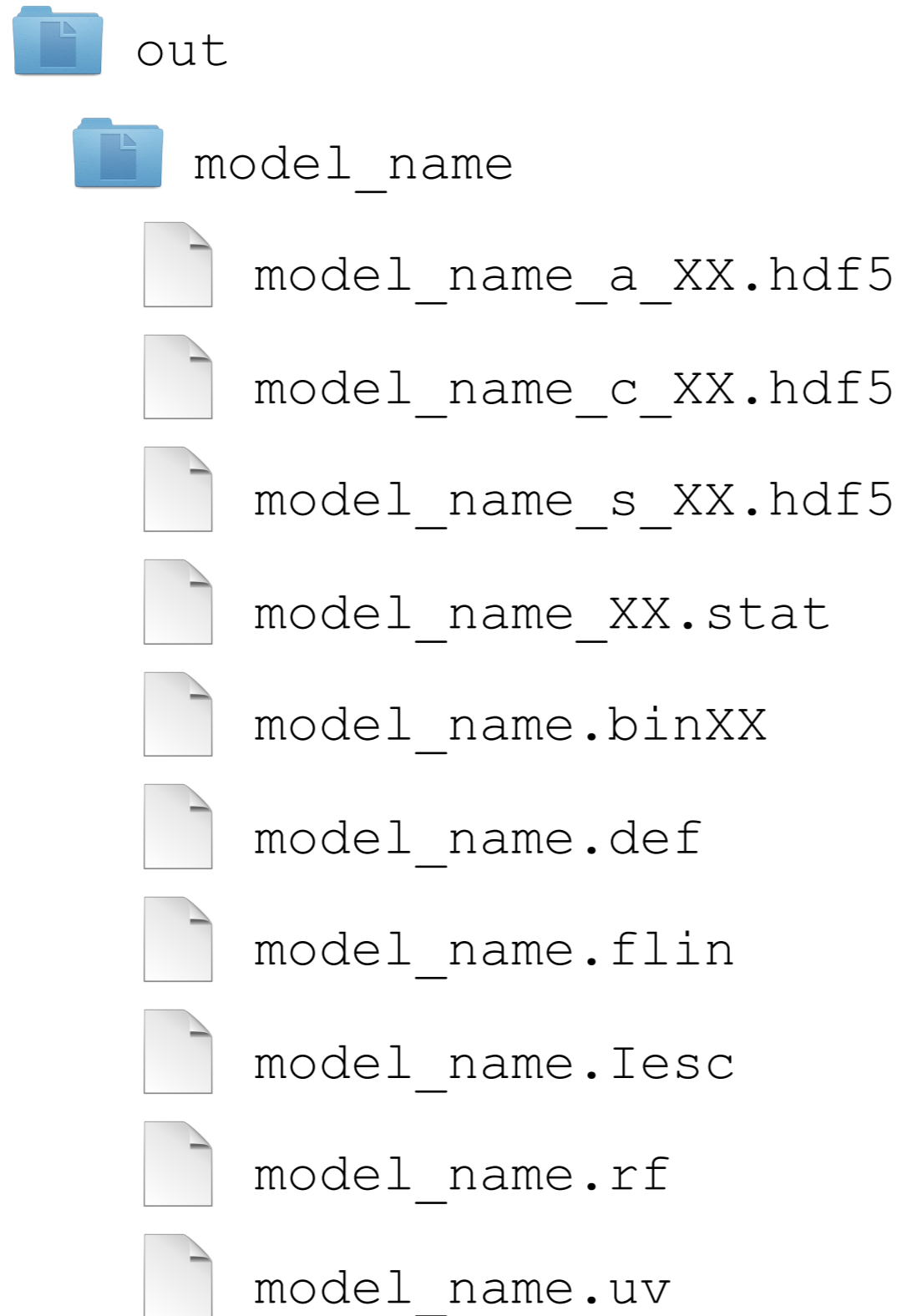
- with another input file `other_input.in`

```
$ ./PDR ../data/other_input.in
```

## Outline

- introduction on PDRs
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- user guide first steps
- examples

# Outputs



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

# Outputs



out



model\_name



model\_name\_a\_XX.hdf5



analysis output



model\_name\_c\_XX.hdf5



chemical output



model\_name\_s\_XX.hdf5



standard output



model\_name\_XX.stat



model\_name.binXX



model\_name.def



log file



model\_name.flin



model\_name.Iesc



model\_name.rf



model\_name.uv

## Outline

- introduction on PDRs
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- user guide first steps
- examples



# Outputs



out



model\_name



model\_name\_a\_XX.hdf5



analysis output



model\_name\_c\_XX.hdf5



chemical output



model\_name\_s\_XX.hdf5



standard output



model\_name\_XX.stat



model\_name.binXX



old binary output



model\_name.def



log file



model\_name.flin



model\_name.Iesc



model\_name.rf



model\_name.uv

## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

# Outputs



out



model\_name



model\_name\_a\_XX.hdf5



analysis output



model\_name\_c\_XX.hdf5



chemical output



model\_name\_s\_XX.hdf5



standard output



model\_name\_XX.stat



model\_name.binXX



old binary output



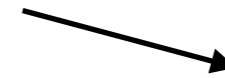
model\_name.def



log file



model\_name.flin



model\_name.Iesc



radiation field



model\_name.rf



related quantities



model\_name.uv



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

# Outputs



out



model\_name



model\_name\_a\_XX.hdf5



analysis output



model\_name\_c\_XX.hdf5



chemical output



model\_name\_s\_XX.hdf5



**standard output**



model\_name\_XX.stat



model\_name.binXX



old binary output



model\_name.def



**log file**



model\_name.flin



model\_name.Iesc



radiation field



model\_name.rf



related quantities



model\_name.uv



## Outline

- introduction on PDRs
- assumptions
- user guide first steps
- examples

## Outline

- introduction on PDRs
- assumptions
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Tools (in `AnalysisTools/` directory)

- read HDF5 files (in `extractor/`)

```
$ python2.7 extractor.py
```

- analyse chemistry (in `ChemistryAnalyser/`)

```
$ python2.7 server/server.py
```

launch `visualizer/index.html` with a web browser

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- introduction on PDRs
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- user guide first steps
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Method

- always check
  - ✓ convergence
  - ✓ temperature, density & ionization profiles
  - ✓ H / H<sub>2</sub> & C<sup>+</sup> / C / CO transitions
- understand local before integrated quantities



## Natural variables for radiative transfer

$$\frac{\partial I_\lambda}{\partial s} = -(\kappa_\lambda + \sigma_\lambda)I_\lambda + \text{scattering} + \text{emission}$$

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$$\frac{\partial I_\lambda}{\partial s} = -(\kappa_\lambda + \sigma_\lambda)I_\lambda + \text{scattering} + \text{emission}$$

absorption & scattering  
coefficients

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## Natural variables for radiative transfer

$$\frac{\partial I_\lambda}{\partial s} = -(\kappa_\lambda + \sigma_\lambda)I_\lambda + \text{scattering} + \text{emission}$$

- optical depth
- extinction
- visible extinction
- H column density

$$d\tau_\lambda = (\kappa_\lambda + \sigma_\lambda)ds$$

$$A_\lambda = 2.5 \log_{10}(e) \tau_\lambda$$

$$A_\lambda = A_V \left( 1 + \frac{k(\lambda - V)}{R_V} \right)$$

$$N_H = C_D / R_V \times A_V$$

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extinction curve

- optical depth
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- visible extinction
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$A_V / E_{B-V}$

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